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I. Existence theorems

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On lattice points in *n*-dimensional star bodies

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Let $F(X) = F(x_1, ..., x_n)$ be a continuous non-negative function of X satisfying F(tX)

=|t|F(X) for all real numbers t. The set K in n-dimensional Euclidean space R_n defined by $F(X) \leq 1$ is called a star body. The author studies the lattices Λ in R_n which are of minimum determinant and have no point except (0, ..., 0) inside K. He investigates how

many points of such lattices lie on, or near to, the boundary of K, and considers in detail the case when K admits an infinite group of linear transformations into itself.

Introduction

Let
$$K$$
 be an arbitrary bounded or unbounded point set in the n -dimensional Euclidean space R of all points

space R_n of all points $X = (x_1, x_2, ..., x_n)$ $(x_1, x_2, ..., x_n \text{ real numbers}).$

A point lattice
$$\Lambda$$
,

$$x_h = \sum_{k=1}^{n} a_{hk} u_k \quad (h = 1, 2, ..., n, u_1, u_2, ..., u_n = 0, \pm 1, \pm 2, ...),$$

 $d(\Lambda) = \left| \left| a_{hk} \right|_{h,k=1,2,\dots,n} \right|$ in R_n of determinant is called K-admissible if no point P of Λ , except possibly the origin O = (0, 0, ..., 0),

is an inner point of K. (P is an inner point of K if there is an n-dimensional sphere with centre at P and contained in K.) The minimum determinant $\Delta(K)$ of K is defined as the lower bound of $d(\Lambda)$ extended over all K-admissible lattices. This function $\Delta(K)$ depends on K in a very complicated way and is, in general, not a continuous function of K. A K-admissible lattice Λ such that $d(\Lambda) = \Delta(K)$ is called

a critical lattice of K; such critical lattices exist, for instance, if K contains O as an

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Minkowski proved in his classical theorem that if K is a convex body with centre at O, then $2^n \Delta(K) \geqslant V(K)$.

where V(K) is the volume of K. He further gave a finite algorism for obtaining $\Delta(K)$ and the critical lattices of K if K is such a convex body and n=2 or n=3, or if K is of a certain type with n=4 (Minkowski 1907, 1911). Minkowski also considered another more general class of point sets, the star

$$F(X) \leqslant 1$$
,

bodies (Strahlenkörper). These are point sets defined by an inequality

where
$$F(X) = F(x_1, ..., x_n)$$
 is a continuous function of X such that $F(X) \ge 0$ for all points X ,

inner point and has at least one admissible lattice.

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$$F(tx_1, ..., tx_n) = |t| F(x_1, ..., x_n)$$
 for real t .

The functional equation implies that K is symmetrical in O. This restriction is not

made by Minkowski, but is in no way essential. He found (1911) for such point

sets that

 $2\zeta(n)\Delta(K) \leq V(K)$,

but his proof was never published. Recently, Hlawka (1943) gave a very ingenious

proof based on the theory of multiple integrals, and I found a geometrical proof

(Mahler 1944) for a slightly less exact inequality.* New progress was made in the years from 1938 onwards when important special

examples of star bodies in two or three dimensions were investigated by Davenport

(1938, 1939 and 1944) and Mordell (1942, 1943, 1944, and the general method 1945). In 1941 Mordell discovered a method for dealing with a certain important class of such problems. This work led me to ask myself whether Minkowski's method of evaluating $\Delta(K)$ when K is convex (Minkowski 1907, 1911) could be extended to

arbitrary bounded star bodies. I succeeded in answering this question in the affirmative, and found an algorism for the evaluation of $\Delta(K)$ if K is two-dimensional

and bounded; and I applied this method to a few special cases. In the present paper, the aim is not to consider further special examples of star

bodies, but rather to lay the foundations of a general theory of bounded or unbounded *n*-dimensional star bodies and their critical lattices.

In this first part, I begin by proving that if the star body K, $F(X) \leq 1$,

* Addition, May 1946. A beautiful new proof of the Minkowski-Hlawka theorem was recently given by C. L. Siegel, Ann. Math. 46 (1945), 340-347.

almost obvious; an example is constructed in which this lower bound is attained. If K is not bounded, then Λ need not have a single point on C, as is also proved by means of an example. It is then easily proved that to every $\epsilon > 0$ there is at least one point P of Λ such that

The points of such a critical lattice Λ on, or in the neighbourhood of, the boundary C of K are next studied. If K is bounded, then at least 2n points of A lie on C, as is

$$1\leqslant F(P)<1+\varepsilon;$$
 however, it remains an open question whether there are always n independent points

of Λ with this property. From § 14 onwards, unbounded star bodies are considered with an infinite group Γ of linear transformations into themselves; many of the most interesting lattice-

$$\Gamma$$
 of linear transformations into themselves; many of the most interesting lattice-point problems are of this type. Three different assumptions about Γ are made and applied to the study of the critical lattices. Then three general classes of star bodies

applied to the study of the critical lattices. Then three general classes of star bodies are found with the following three properties respectively: (a) At least one critical lattice of
$$K$$
 has a point on C (theorem 21). (b) For every $\epsilon > 0$, every critical lattice A of K contains an infinity of points P satisfying

 $1 \leq F(P) < 1 + \epsilon$

(theorem 23). (c) For every
$$\epsilon > 0$$
, every critical lattice Λ of K contains n independent points P_1, \ldots, P_n satisfying

$$1\leqslant F(P_g)<1+\epsilon\quad (g=1,2,...,n)$$
 (theorem 25). The simplest example of an n -dimensional star body with all three

properties (a), (b), (c) is that defined by the inequality
$$\big|\,x_1x_2\dots x_n\,\big|\leqslant 1.$$

contain, or do not contain, smaller star bodies K' such that

 $\Delta(K') = \Delta(K).$

In the second part of this paper which is appearing in the Proc. Royal Acad. Amsterdam, I intend to study certain types of star bodies K according as they

1. Notation

If $x_1, x_2, ..., x_n$ $(n \ge 2)$ are real numbers, then

 $X = (x_1, x_2, \dots, x_n)$

(1.1)

(1.2)

$$X=(x_1,x_2)$$

The following notation is used in this paper:

is the point in n-dimensional Euclidean space R_n with rectangular co-ordinates

 $x_1, x_2, ..., x_n$. The non-negative number $|X| = +(x_1^2 + x_2^2 + \dots + x_n^2)^{\frac{1}{2}}$

154 K. Mahler is called the distance of X from the origin O = (0, 0, ..., 0). If $X_1 = (x_1^{(1)}, x_2^{(1)}, \dots, x_n^{(1)}), \dots, X_r = (x_1^{(r)}, x_2^{(r)}, \dots, x_n^{(r)})$ are any points in R_n , and $\lambda_1, ..., \lambda_r$ are any real numbers, then $\lambda_1 X_1 + \ldots + \lambda_r X_r$ is written for the point $(\lambda_1 x_1^{(1)} + \ldots + \lambda_r x_1^{(r)}, \lambda_1 x_2^{(1)} + \ldots + \lambda_r x_2^{(r)}, \ldots, \lambda_1 x_n^{(1)} + \ldots + \lambda_r x_n^{(r)})$ The determinant of n points $X_1 = (x_1^{(1)}, x_2^{(1)}, \dots, x_n^{(1)}), \dots, X_n = (x_1^{(n)}, x_2^{(n)}, \dots, x_n^{(n)})$ (1.4) $\{X_1,X_2,\ldots,X_n\} = \left| \begin{array}{cccc} x_1^{\alpha'} & x_2^{\alpha'} & \ldots & x_n^{\alpha'} \\ x_1^{(2)} & x_2^{(2)} & \ldots & x_n^{(2)} \\ \vdots & \vdots & & \vdots \\ x^{(n)} & x_n^{(n)} & \ldots & x_n^{(n)} \end{array} \right|.$ is denoted by (1.5)The points are called independent, if this determinant does not vanish. The set Λ of all points

 $X = u_1 X_1 + ... + u_n X_n$, where $u_1, ..., u_n = 0, \pm 1, \pm 2, \pm 3, ...$ is called a lattice if its determinant

 $d(\Lambda) = |\{X_1, X_2, ..., X_n\}|$ (1.6)is not zero; then $X_1, X_2, ..., X_n$ are said to form a basis of Λ . Any n points $Y_1, Y_2, ..., Y_n$

of Λ form a basis of this lattice if and only if $\{Y_1, Y_2, \dots, Y_n\} = \pm d(\Lambda).$ (1.7)If P, Q, R, \dots are points of Λ , then $\Lambda - [P, Q, R, \dots]$ denotes the set of all points

of Λ different from P, Q, R, \dots

2. The reduced basis of a lattice

Theorem 1. There exists a constant $\gamma_n > 0$ depending only on the dimension n

THEOREM 1. There exists a constant
$$\gamma_n > 0$$
 depending only on the dimension of R_n , with the following property: every lattice A in R_n has a reduced basis, i.e. a bas

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if
$$R_n$$
, with the following property: every lattice Λ in R_n has a reduced basis, i.e. a bas Y_1, Y_2, \ldots, Y_n for which

 $Y_1, Y_2, ..., Y_n$ for which

$$|Y_1,Y_2,\ldots,Y_n|$$
 for which
$$|Y_1|\,|Y_2|\ldots|\,Y_n\,|\leqslant \gamma_n\,d(A). \tag{2}$$

 $(2 \cdot 1)$

$$|Y_1| |Y_2| \dots |Y_n| \leqslant \gamma_n d(\Lambda). \tag{2}$$
Proof Let $Y = (x^{(1)}, x^{(1)}, x$

Proof. Let $X_1 = (x_1^{(1)}, x_2^{(1)}, \dots, x_n^{(1)}), \dots, X_n = (x_1^{(n)}, x_2^{(n)}, \dots, x_n^{(n)})$ $(2 \cdot 2)$

be any basis of Λ . Then

e any basis of
$$\Lambda$$
. Then

 $\varPhi(u_1, \ldots, u_n) = \sum\limits_{g=1}^n (x_g^{(1)} u_1 + \ldots + x_g^{(n)} u_n)^2 = \big| \; u_1 X_1 + \ldots + u_n \, X_n \, \big|^2$

 $(2 \cdot 3)$

is a positive definite quadratic form of discriminant

 $d(\Lambda)^2 = \{X_1, X_2, \dots, X_n\}^2.$ $(2 \cdot 4)$

(2.10)

(3.1)

(3.2)

(3.3)

(3.4)

 $\{Y_1, \dots, Y_n\} = \{X_1, \dots, X_n\} = +d(A).$ Moreover,

 $\Psi(1,0,...,0) = |Y_1|^2, \ \Psi(0,1,...,0) = |Y_2|^2, \ ..., \ \Psi(0,0,...,1) = |Y_n|^2,$

whence the assertion. Theorem I may also be proved by the reduction method of Hermite (1905),

which has the advantage that the proof of the product formula for the Ψ 's is of an elementary character.

The convergence theorem Definition 1. An infinite sequence of lattices

 A_1, A_2, A_3, \dots

is called bounded, if there exist two positive numbers c_1 , c_2 such that

 $d(\Lambda_r) \leq c_1$ for $r = 1, 2, 3, \dots$

 $|X| \geqslant c_2$ for all points $X \neq O$ of A_r , when r = 1, 2, 3, ...

Definition 2. An infinite sequence of lattices

 $\Lambda_1, \Lambda_2, \Lambda_2, \dots$

is said to converge, and to have as its limit the lattice Λ , if there exist reduced bases

 $Y_1^{(r)}, Y_2^{(r)}, \dots, Y_n^{(r)} \text{ of } \Lambda_r \text{ for } r = 1, 2, 3, \dots$

and a basis such that

 Y_1, Y_2, \ldots, Y_n of Λ , $\lim_{r \to \infty} \left| \begin{array}{c} Y_g^{(r)} - Y_g \end{array} \right| = 0, \quad where \quad g = 1, 2, ..., n.$

This definition implies that the points of A_r in any finite region independent of rtend to the points of Λ , as r tends to infinity. From these two definitions is derived the following theorem which is fundamental for the study of star bodies:

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Theorem 2. Every bounded infinite sequence of lattices contains a convergent infinite subsequence. *Proof.* Let $\Lambda_1, \Lambda_2, \Lambda_3, \ldots$ be any bounded sequence, and let $Y_1^{(r)}, Y_2^{(r)}, \ldots, Y_n^{(r)}$ be

a reduced basis of Λ_r for r = 1, 2, 3, ..., then from definition 1, $d(\Lambda_r) \leq c_1, \quad |Y_q^{(r)}| \geq c_2, \quad \text{where} \quad g = 1, 2, ..., n \text{ and } r = 1, 2, 3, ...,$ (3.5)and from theorem 1, (3.6)

$$|Y_1^{(r)}| |Y_2^{(r)}| \dots |Y_n^{(r)}| \leqslant \gamma_n d(\Lambda_r), \quad \text{where} \quad r=1,2,3,\dots, \tag{3.6}$$
 hence
$$|Y_g^{(r)}| \leqslant \gamma_n c_1 c_2^{-(n-1)}, \quad \text{where} \quad g=1,2,\dots,n \text{ and } r=1,2,3,\dots \tag{3.7}$$
 All co-ordinates of the basis points $Y_g^{(r)}$ $(g=1,2,\dots,n; r=1,2,3,\dots)$ are therefore bounded, and so there exists an infinite sequence of indices

 $r_1, r_2, r_3, \ldots,$ Y_1, Y_2, \ldots, Y_n

and a set of
$$n$$
 points $Y_1, Y_2, ..., Y_n$, such that $\lim_{k \to \infty} |Y_g^{(r_k)} - Y_g| = 0$, where $g = 1, 2, ..., n$, (3.8)

such that (3.9)

 $\lim_{k\to\infty}d(A_{r_k})=\lim_{k\to\infty}\big|\left\{Y_1^{(r_k)},\,Y_2^{(r_k)},\,\ldots,\,Y_n^{(r_k)}\right\}\big|=\big|\left\{Y_1,Y_2,\,\ldots,Y_n\right\}\big|.$ whence $\gamma_n d(\Lambda_{r_k}) \geqslant |Y_1^{(r_k)}| |Y_2^{(r_k)}| \dots |Y_n^{(r_k)}| \geqslant c_2^n,$ (3.10)Further, from $d(\Lambda_{r_n}) \geqslant \gamma_n^{-1} c_2^n$ and

(3.11) $|\{Y_1, Y_2, \dots, Y_n\}| \ge \gamma_n^{-1} c_2^n > 0,$ (3.12)it is deduced that

and so the lattice Λ of basis $Y_1, Y_2, ..., Y_n$ satisfies the assertion. 4. Distance functions and star bodies Definition 3. A function

 $F(X) = F(x_1, x_2, ..., x_n)$ (4.1)of the point $X = (x_1, x_2, ..., x_n)$ in R_n is called a distance function if it satisfies the

following conditions: (a) $F(X) \ge 0$ for all points, and F(X) > 0 for at least one point;

(b) F(tX) = |t| F(X) for all points X and all real numbers t; hence F(-X) = F(X) and F(O) = 0;

(c) F(X) is a continuous function of X.

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(5.1)

that all points of the sphere $|X| \leq \rho$ belong to K.

It is evident that a star body K has the following properties: (A) If X belongs to K, then tX, where $-1 \le t \le 1$, also belongs to K. (B) The limit point of a convergent sequence of points of K also belongs to K.

(C) The origin O is an inner point of K; i.e. there exists a positive number ρ such

Definition 4. The set K of all points X satisfying $F(X) \leq 1$ is called the star body

For since F(X) is continuous, it assumes on the sphere |X| = 1 a maximum value, say $1/\rho$. Then $F(X) \mid X \mid^{-1} \le 1/\rho$ for all $X \ne 0$, whence $F(X) \le 1$, if X is a point of the sphere $|X| \leq \rho$.

Theorem 3. The star body K is bounded if and only if F(X) > 0 for all points $X \neq 0$.

Proof. As a continuous function, F(X) assumes on the sphere |X| = 1 a minimum, say μ . If $\mu = 0$, then F(X) vanishes at a point $X \neq 0$, and so it vanishes at all points of the line through O and X; hence K is not bounded. If, however, $\mu = 1/P > 0$, then $F(X) \mid X \mid^{-1} \ge 1/P$ for all $X \ne 0$, hence $\mid X \mid \le P$ if $F(X) \le 1$, and so K is bounded.

5. The two types of star bodies

Definition 5. The lattice Λ is called K-admissible if $\Lambda - [O]$ contains no inner points of K.

Definition 6. The star body K is called of the finite type if there exists at least one K-admissible lattice; it is called of the infinite type if no such lattice exists. Theorem 4. Every bounded star body is of the finite type.

Proof. Let P > 0 be a number such that $|X| \leq P$ for all points of K, and denote by Λ the lattice of basis

 $X_1 = (P, 0, ..., 0), X_2 = (0, P, ..., 0), ..., X_n = (0, 0, ..., P).$

Then $|X| \ge P$ for all points $X \ne O$ of Λ ; hence Λ is K-admissible. THEOREM 5. Unbounded star bodies exist of the finite type, and also of the infinite

type.

Proof. (1) The star body K of distance function

 $F(X) = |x_1 x_2 \dots x_n|^{1/n}$

(5.2)is not bounded. To show that K is of the finite type, denote by \Re any totally real

algebraic field of degree n, by $\omega_1^{(g)}, \omega_2^{(g)}, \ldots, \omega_n^{(g)}, \text{ where } g = 1, 2, \ldots, n,$

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conjugate inte	egral bases of the n fields $\Re^{(1)}, \Re^{(2)}, \dots, \Re^{(n)}$ conjugate to \Re basis	, and by Λ
	$X_h = (\omega_h^{(1)}, \omega_h^{(2)}, \dots, \omega_h^{(n)}), \text{ where } h = 1, 2, \dots, n.$	(5.3)
hence $F(X) \geqslant$	For the sign, $F(X)$ is the norm of an integer $\alpha \neq 0$ in \Re if X lies 1 for all lattice points $X \neq O$. It is body of distance function	$\sin A - [O];$
	$F(X) = x_1^2 x_2 \dots x_n ^{1/(n+1)}$	(5.4)
denote by t_1, t_2	bounded, but it is of the infinite type. For let Λ be any t_0, \ldots, t_n , n positive numbers of product $d(\Lambda)$. By Minkowsk as, there exists a point $X = (x_1, x_2, \ldots, x_n) \neq 0$ of Λ such that	i's theorem
	$ x_1 \le t_1, x_2 \le t_2,, x_n \le t_n,$	(5.5)
hence	$0 < F(X) \le \{t_1 d(A)\}^{1/(n+1)}.$	(5.6)
not K -admissi	ed now that $t_1 < d(A)^{-1}$, then X is an inner point of K. Thible. Therefore, all star bodies considered are from now on a	
be of the finite	e type.	
	6. The determinant of a star body	
`	$(7) \leq 1$, be a star body of the finite type. By definition 6, the same lattices is not empty. Hence the lower bound	ne set $\Lambda(K)$
	$\Delta(K) = 1.b. d(\Lambda)$	(6.1)
	r all elements of $A(K)$, exists; $\Delta(K)$ is called the determines K of the infinite type, put $\Delta(K) = \infty$.	nant of K .
Theorem 6.	. The determinant of a star body is positive.	
<i>Proof.</i> By thence also the	the property (C) of a star body (§ 4), K contains the sphe cube	re $ X \leq \rho$,
	$\max(x_1 , x_2 ,, x_n) \le \rho n^{-\frac{1}{2}}.$	(6.2)
By Minkowski	i's theorem on linear forms, every lattice of determinant	
	$d(\Lambda) < \rho^n n^{-\frac{1}{2}n}$	
	nner point $X \neq O$ of this cube, i.e. of K , and so such a lattice Hence, for every K -admissible lattice Λ ,	e cannot be
	$d(\Lambda) \geqslant \rho^n n^{-\frac{1}{2}n},$	(6.3)
whence	$\Delta(K) \geqslant \rho^n n^{-\frac{1}{2}n} > 0.$	(6.4)
Theorem 7.	. If the star body H is contained in the star body K , then	
	$\Delta(H) \leqslant \Delta(K)$.	(6.5)
Proof. Even	ry K -admissible lattice is also H -admissible.	

(7.1)

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and $d(\Lambda) = \Delta(K)$.

Theorem 8. Every star body of the finite type possesses at least one critical lattice. *Proof.* From the definition of $\Delta(K)$, there exists an infinite sequence of K-admis-

sible lattices

The following theorem is fundamental for the theory:

$$A_1, A_2, A_3, \dots,$$

not necessarily all different, such that

$$\lim_{r \to \infty} d(\Lambda_r) = \Delta(K);$$

it may be assumed further, without loss of generality, that

ther, without loss of generality, that
$$d(\Lambda_r) \leq 2\Delta(K)$$
, where $r = 1, 2, 3, \ldots$

$$d(A_r) \leqslant 2\Delta(K)$$
, where $r=1$, Moreover, since the sphere $|X| \leqslant \rho$ is contained in K , $|X| \geqslant \rho$ for all points $X \neq O$ of A_r , when

$$|X| \geqslant \rho$$
 for all points $X \neq 0$ of Λ_r ,
From (7·2) and (7·3) the sequence $\{\Lambda_r\}$ is boun
it contains a convergent infinite subsequence

 $\lim_{k \to \infty} |Y_g^{(r_k)} - Y_g| = 0$, where g = 1, 2, ..., n,

 $Y = u_1 Y_1 + \ldots + u_n Y_n + O$, where u_1, \ldots, u_n are integers

 $\lim_{t \to \infty} |Y^{(r_k)} - Y| = 0.$

Hence $Y^{(r_k)} \neq O$ for sufficiently large k, and so $F(Y^{(r_k)}) \geqslant 1$ since A_{r_k} is K-admissible,

 $F(Y) = \lim_{k \to \infty} F(Y^{(r_k)}) \geqslant 1.$

oreover, since the sphere
$$|X| \leq \rho$$
 is contained in K , $|X| \geq \rho$ for all points $X \neq 0$ of A_r , where $r = 1, 2, 3, \ldots$ (7·3) From (7·2) and (7·3) the sequence $\{A_r\}$ is bounded, and hence, from theorem 2, contains a convergent infinite subsequence

$$r = 1, 2$$
 and hence

$$(7.2)$$

$$(7.3)$$
heorem 2,

say of limit
$$A$$
. Denote by $Y_1^{(r_k)}, Y_2^{(r_k)}, ..., Y_n^{(r_k)}$ a reduced basis of A_{r_k} , by $Y_1, Y_2, ..., Y_n$ a basis of A , taken such that
$$\lim_{k \to \infty} |Y_g^{(r_k)} - Y_g| = 0, \quad \text{where} \quad g = 1, 2, ..., n, \tag{7.4}$$
 hence
$$d(A) = |\{Y_1, Y_2, ..., Y_n\}| = \lim_{k \to \infty} |\{Y_1^{(r_k)}, Y_2^{(r_k)}, ..., Y_n^{(r_k)}\}| = \lim_{k \to \infty} d(A_{r_k}) = \Delta(K). \tag{7.5}$$
 Let further
$$Y = u_1 Y_1 + ... + u_n Y_n \neq 0, \quad \text{where} \quad u_1, ..., u_n \text{ are integers} \tag{7.6}$$

(7.4) $d(A) = \left| \{Y_1, Y_2, ..., Y_n\} \right| = \lim_{k \to \infty} \left| \{Y_1^{(r_k)}, Y_2^{(r_k)}, ..., Y_n^{(r_k)}\} \right| = \lim_{k \to \infty} d(A_{r_k}) = \Delta(K). \quad (7.5)$

(7.7)

(7.8)

(7.9)

hence

whence

$$Y^{(r_k)}=u_1\,Y_1^{(r_k)}+\ldots+u_n\,Y_n^{(r_k)},\quad\text{where}\quad k=1,2,3,\ldots;$$
 then
$$\lim \|Y^{(r_k)}-Y\|=0.$$

a basis of Λ , taken such that

be any point of
$$\varLambda$$
, and put
$$Y^{(r_k)} = u_1 \, Y^{(r_k)}$$

From (7.5) and (7.9), Λ satisfies the assertion.

K. Mahler 1608. The continuity of $\Delta(K)$ If $K: F(X) \leq 1$, is any star body, and if t is a positive number, then we denote by tK the star body of distance function $t^{-1}F(X)$, i.e. the set of all points X for which $F(X) \leq t$. From homogeneity, it is evident that $\Delta(tK) = t^n \Delta(K).$ (8.1)The set of all points X in K for which $|X| \leq t$ is further denoted by K^t . THEOREM 9. Let K, K_1, K_2, \ldots be an infinity of star bodies of the finite type, satisfying the following conditions: (a) To every $\epsilon > 0$, there is a positive integer $N(\epsilon)$ such that K_r is contained in $(1 + \epsilon)$ K_r if $r \geqslant N(\epsilon)$. (b) To every t > 0 and every $\epsilon > 0$, there is a positive integer $N(t, \epsilon)$ such that K^t is contained in $(1+\epsilon) K_r$ if $r \ge N(t, \epsilon)$. $\lim_{r\to\infty} \Delta(K_r) = \Delta(K).$ (8.2)Then*Proof.* From (a), by theorem 7, $\Delta(K_n) \leq \Delta((1+\epsilon)K) = (1+\epsilon)^n \Delta(K),$ (8.3) $\limsup \Delta(K_r) \leqslant \Delta(K).$ (8.4)whence for $\epsilon \to 0$,

It will now be shown that also $\lim\inf \Delta(K_r) \geqslant \Delta(K).$ (8.5)

Let this inequality be false. Then there exists an infinite sequence of indices r_1, r_2, r_3, \dots not smaller than $N(\rho, 1)$ such that

 $\varDelta(K_{r_k}) \leqslant 2 \varDelta(K), \quad \text{and} \quad \lim_{k \to \infty} \varDelta(K_{r_k}) < \varDelta(K).$ (8.6)Denote by A_{r_k} a critical lattice of K_{r_k} ; therefore

 $d(\Lambda_{r_k}) \leqslant 2\Delta(K)$. (8.7)Then from (b) above, on taking $t = \rho$, $\epsilon = 1$, the star body $2K_{r_k}$ contains K^{ρ} , i.e. the

sphere $\mid X \mid \leq \rho$; hence K_{r_k} contains the sphere $\mid X \mid \leq \frac{1}{2}\rho$. Since Λ_{r_k} is K_{r_k} -admissible,

this implies that $|X| \geqslant \frac{1}{2}\rho$ for all points $X \neq 0$ of $\Lambda_{r_{\nu}}$.

It is clear from this and (8.7) that the sequence of lattices $\{A_{r_k}\}$ is bounded.

Therefore, from theorem 2, this sequence contains a convergent infinite subsequence

 $\Lambda^{(1)} = \Lambda_{r_{k_2}}, \ \Lambda^{(2)} = \Lambda_{r_{k_2}}, \ \Lambda^{(3)} = \Lambda_{r_{k_2}}, \ \dots,$

(8.8)of limit Λ , say. For shortness write

 $K^{(1)} = K_{r_{k_2}}, \ K^{(2)} = K_{r_{k_2}}, \ K^{(3)} = K_{r_{k_2}}, \ \dots,$

(8.9)

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(8.11)

(8.14)

(8.15)

(8.16)

(8.17)

(8.18)

(8.19)

ΙI

 $d(\Lambda) = \lim_{k \to \infty} \Delta(K_{r_k}) < \Delta(K).$ This means that Λ is not K-admissible; hence Λ contains at least one point $Y \neq 0$ which is an inner point of K. Denote now by $Y_1^{(l)}, Y_2^{(l)}, ..., Y_n^{(l)}$ a reduced basis of $A^{(l)}$, and by $Y_1, Y_2, ..., Y_n$ a basis

of Λ taken such that $\lim_{I \to \infty} |Y_g^{(l)} - Y_g| = 0$, where g = 1, 2, ..., n; (8.12)then Y can be written as $Y = u_1 Y_1 + ... + u_n Y_n$ (8.13)

as
$$u_1, \ldots, u_n$$
 not all zero. Now put
$$Y^{(l)} = u_1 Y_1^{(l)} + \ldots + u_n Y_n^{(l)},$$

 $F(Y^{(l)}) \leq \frac{1}{1 + 2c}$

and so $(1+2\epsilon) Y^{(l)} \neq 0$ belongs to K. This implies, from (8·16), that $(1+2\epsilon) Y^{(l)}$ is a point of K^t , where $t = 2(1+2\epsilon) | Y |$. Hence, from (b) above, the point $(1+2\epsilon) Y^{(b)}$ belongs to $(1+\epsilon) K^{(l)}$ if l is sufficiently large. This implies that $Y^{(l)}$ is a point of $\frac{1+\epsilon}{1+2\epsilon}K^{(l)}$ and so is an inner point of $K^{(l)}$. However, this is impossible since $Y^{(l)} \neq O$

THEOREM 10. Let $K: F(X) \leq 1$ be a star body of the finite type, G(X) an arbitrary

 $\lim_{t\to\infty} \Delta(K_t) = \Delta(K).$

Proof. It is evident from definition 3 that $F_{\ell}(X)$ is a distance function. Since $F_t(X) \ge F(X)$ for all X and t, K_t is contained in K and so is a star body of the finite

 $Y^{(l)} = u_1 Y_1^{(l)} + \ldots + u_n Y_n^{(l)},$

en
$$Y^{(l)}$$
 belongs to $A^{(l)},$
$$Y^{(l)} \neq O, \quad \text{and} \quad \lim \mid Y^{(l)} - Y \mid = 0,$$

whence, for sufficiently large indices l,

hence, for sufficiently large indices
$$l$$
,

Since Y is an inner point of K and different from O, there is an $\epsilon > 0$ such that

hence, for sufficiently large indices
$$l$$
, $\mid Y^{(l)} \mid \leq 2 \mid Y \mid$. Since Y is an inner point of K and different from $F(Y) \leq \frac{1}{1+3\epsilon}$,

hence, if l is sufficiently large, from (8·15) it follows that

distance function, and t a positive parameter. Then the star body

$$Y^{(l)}\! \neq\! O, \quad \text{and}$$
 nence, for sufficiently large indices $l,$

en
$$Y^{(l)}$$
 belongs to $A^{(l)},$
$$Y^{(l)} \! = \! O, \quad \text{an}$$
 nence, for sufficiently large indices

$$Y^{(l)} \pm O$$
, and hence, for sufficiently large indices l , $\mid Y^{(l)} \mid$

$$Y^{(l)} \neq O$$
, and hence, for sufficiently large indices l , $\mid Y^{(l)} \mid$

 $Y^{(l)} \neq O$, and $\lim_{l \to \infty} |Y^{(l)} - Y| = 0$, $|Y^{(l)}| \leq 2 |Y|$.

with integral coefficients u_1, \ldots, u_n not all zero. Now put then $Y^{(l)}$ belongs to $A^{(l)}$,

and so

 $K_t: F_t(X) \le 1$, where $F_t(X) = \max(F(X), t^{-1}G(X))$, is also of the finite type, and further

and since $\Lambda^{(l)}$ is a critical lattice of $K^{(l)}$.

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type. Further, since the set $H: G(X) \leq 1$ is a star body, there exists a number $\tau > 0$ such that H contains the whole sphere $|X| \leq \tau$. The sphere $|X| \leq \tau t$ is then contained in tH, and so $K^{\tau t}$, which is a subset of this sphere, is contained in K_t . The hypothesis of theorem 9 is therefore satisfied, and so

 $\Delta(K) = \lim_{r \to \infty} \Delta(K_{t_r})$

for every sequence of positive numbers t_1, t_2, t_3, \ldots of limit infinity. This proves the

(8.20)

(8.21)

(10.1)

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Remark. The results of this paragraph remain true when $\Delta(K) = \infty$.

9. Lattice points on the boundary of a bounded star body

and generate L, while $Y_{f+1}, ..., Y_n$ lie outside L. Let $\epsilon > 0$ be sufficiently small and

Originally (Mahler 1943), I used this formula as the definition of $\Delta(K)$ for unbounded star bodies, so reducing the problem to one for the bounded case.

 $\Delta(K) = \lim_{t \to \infty} \Delta(K^t).$

The last theorem, for G(X) = |X|, shows that

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assertion.

Theorem 11. If K is a bounded star body, then every critical lattice of K has n independent points on the boundary C of K. *Proof.* Let Λ be a K-admissible lattice which does not contain n independent

points on C. Then denote by Π the set of all points of Λ on C, and by L the linear space of lowest dimension $f(0 \le f \le n-1)$ containing Π . By Minkowski's method of adaptation of lattices, a basis $Y_1, ..., Y_n$ of Λ can be found such that $Y_1, ..., Y_t$ lie in

denote by A^* the lattice of basis $Y_1, ..., Y_f, (1-\epsilon)Y_{f+1}, ..., (1-\epsilon)Y_n$. This lattice is K-admissible since O and the elements of Π are its only points belonging to K. Since $d(\Lambda^*) = (1-\epsilon)^{n-f} d(\Lambda) < d(\Lambda), \Lambda^*$ is of smaller determinant than Λ , and so Λ is not critical. This theorem shows that in the case of a bounded star body K, every critical lattice A has at least 2n points on its boundary C, namely, n independent points P_1, \ldots, P_n

and their images $-P_1, ..., -P_n$ in O. If $\pm P_1, \pm P_2, ..., \pm P_n$ are the only points on C of the lattice Λ , then Λ is called a singular lattice of K; otherwise it is called a regular lattice. The example in the next paragraph shows that

star bodies with singular lattices do exist.

10. An example of a star body with a singular lattice

Theorem 12. There exists a bounded star body with just one critical lattice. Moreover, this lattice is singular.

Proof. Let ϵ be so small a positive constant that $(1-\epsilon)^n > \tfrac{9\,9}{1\,0\,0}, \quad (1-\epsilon)^{n-1}\sqrt[n]{\frac{3}{2}} > 1, \quad \epsilon < n^{-1} \big(\sqrt[n]{(\tfrac{1\,1}{1\,0})} - 1\big)\,,$ $P = tQ + (t-1) \epsilon R$, where $t \ge 1$ and $|R| \le 1$,

describes all points of the unit sphere $|R| \leq 1$, then P lies on or in a sphere centre at tQ and radius $(t-1)\epsilon$; on varying t, we obtain $S_{\epsilon}(Q)$ as the sum set of these spheres. Denote further by A_0 the lattice of all points with integral co-ordinates, i.e. of basis

is a cone with vertex at Q and its open side away from O. For when t is fixed and R

 $P_1 = (1, 0, ..., 0), P_2 = (0, 1, ..., 0), ..., P_n = (0, 0, ..., 1),$ (10.3)and of determinant $d(\Lambda_0) = 1$. (10.4)

The cube
$$W\colon \quad \left|x_1\right|\leqslant \sqrt[n]{\frac{3}{2}}, \ \left|x_2\right|\leqslant \sqrt[n]{\frac{3}{2}}, \ ..., \ \left|x_n\right|\leqslant \sqrt[n]{\frac{3}{2}}$$
 (10·4) contains 3^n points of A_0 , namely, the origin O , the $2n$ points $\pm P_1, \pm P_2, ..., \pm P_n$,

contains
$$3^n$$
 points of A_0 , namely, the origin O , the 2^n and the m points $P'_1, P'_2, ..., P'_m$, where

$$P'_h = (x_1^{(h)}, x_2^{(h)}, \dots, x_n^{(h)}), \quad x_g^{(h)} = 0, 1, \text{ or } -1, \quad \sum_{g=1}^n |x_g^{(h)}| \geqslant 2.$$

Denote by
$$K$$
 the set of all those points of W which are not inner points of one of the cones

tes
$$+P_{r}$$
), where $q=1,2,\ldots,n_{r}$ or SI

$$S_{\epsilon}(\pm P_g)$$
, where $g=1,2,...,n$, or $S_{\epsilon}[(1-\epsilon)\,P_h']$, where $h=1,2,...,m$.

$$S_{\epsilon}(\pm P_g)$$
, where $g=1,2,\ldots,n,$ or $S_{\epsilon}[\cdot]$
Then K is a bounded star body, and the cube

ded star body, and the cube
$$V\colon |x_1| \le 1 - \epsilon, |x_2| \le 1 - \epsilon, ..., |x_n| \le 1 - \epsilon,$$

 $\Delta(K) \geqslant \Delta(V) = (1 - \epsilon)^n > \frac{5}{6}$.

 $\Delta(K) \leq d(\Lambda_0) = 1$,

since, by the construction, Λ_0 is K-admissible. Hence, if Λ is any critical lattice

 $\frac{5}{6} < d(\Lambda) \leq 1$.

 $U_{\boldsymbol{a}} \colon \quad \mid \boldsymbol{x}_{\boldsymbol{a}} \mid \leqslant \sqrt[n]{\frac{3}{2}}, \quad \mid \boldsymbol{x}_{l} \mid \leqslant 1 - e \quad \text{for} \quad l = 1, 2, \ldots, g - 1, g + 1, \ldots, n$

Hence, from Minkowski's theorem on linear forms, at least one point of A - [O] is an inner point of U_g , say the point $P_g^* = (\xi_1^{(g)}, \xi_2^{(g)}, \dots, \xi_n^{(g)})$. This point lies in one of the two cones $S_{\epsilon}(\pm P_g)$, since the other inner points of U_g are also inner points of K. There is no loss of generality in assuming that P_q^* belongs to $S_{\epsilon}(P_q)$ and so may be

 $P_{\sigma}^{*}=t_{\sigma}P_{\sigma}+\left(t_{g}-1\right)\epsilon R_{g},\quad\text{where}\quad t_{g}\geqslant1\text{ and }\left|\;R_{g}\;\right|\leqslant1.$

 $2^{n}(1-\epsilon)^{n-1}\sqrt[n]{\frac{3}{2}} > 2^{n}$.

$$S_{e}[(1-\epsilon) I$$

$$(1-\epsilon) P_h'$$
],

$$(\epsilon) P'_b$$
, where

$$P'_h$$
, where $h = 1$,

Then
$$K$$
 is a bounded star body, and the cube
$$V\colon \quad \big|\, x_1 \,\big| \leqslant 1 - \epsilon, \,\, \big|\, x_2 \,\big| \leqslant 1 - \epsilon, \,\, \ldots, \,\, \big|\, x_n \,\big| \leqslant 1 - \epsilon, \qquad (10 \cdot 6)$$
 obviously is a subset of K . Therefore from theorem 7, Minkowski's theorem on linear

(10.7)

(10.8)

(10.9)

(10.10)

(10.11)

(10.12)II-2

$$\pm P_2, ..., \pm P_n,$$
2. (10.5)

(10.2)

forms, and from $(10\cdot1)$

from $(10 \cdot 1)$ is of volume

Each one of the *n* parallelepipeds

On the other hand

of K, then

written as

K. Mahler Therefore if, say, $R_q = (\eta_1^{(g)}, \eta_2^{(g)}, ..., \eta_n^{(g)})$, then $\xi_a^{(g)} = t_a + (t_a - 1) \, \epsilon \eta_a^{(g)}$ (10.13) $\eta_a^{(g)} \ge -1;$ and (10.14)and since P_q lies in U_q , $\sqrt[n]{\frac{3}{2}} \geqslant \xi_{\alpha}^{(g)} = t_{\alpha} + (t_{\alpha} - 1) \epsilon \eta_{\alpha}^{(g)} \geqslant t_{\alpha} - (t_{\alpha} - 1) \epsilon > (1 - \epsilon) t_{\alpha}$ (10.15) $1 \leq t_g < \frac{\frac{n/\frac{3}{2}}{1-\epsilon}}{1-\epsilon}$, where g = 1, 2, ..., n. whence (10.16)

nence
$$1\leqslant t_g<\frac{\sqrt{2}}{1-\epsilon},$$
 Denote now by D the determinant

and by

whence

and

 $g_1, g_2, ..., g_r$. Hence

 $D = \{P_1^*, P_2^*, \dots, P_n^*\},\$

by
$$E$$
 the unit determinant
$$E=\{P_1,P_2,\ldots,P_n\}=d(A_0)=+1,$$
 and by
$$E(g_1,g_2,\ldots,g_r), \quad \text{where} \quad 1\leqslant r\leqslant n, \ 1\leqslant g_1< g_2<\ldots< g_r< n,$$
 the determinant which is obtained from E if the points P_1,P_2,\ldots,P_n in

the determinant which is obtained from
$$E$$
 if the points $P_{g_1}, P_{g_2}, \ldots, P_{g_r}$ in it are replaced by the points $R_{g_1}, R_{g_2}, \ldots, R_{g_r}$ of the same indices. Obviously $E(g_1, g_2, \ldots, g_r)$ is equal to its cofactor of order r belonging to the rows and columns of indices g_1, g_2, \ldots, g_r . Hence

is equal to its cofactor of order
$$r$$
 belonging to the rows and columns of indices $g_1, g_2, ..., g_r$. Hence $|E(g_1, g_2, ..., g_r)| \le r!$ (10·19) since the moduli of the co-ordinates of $R_{g_1}, R_{g_2}, ..., R_{g_r}$ are not larger than 1, and since a determinant of order r consists of $r!$ terms.

(10.17)

(10.18)

(10.19)

since a determinant of order
$$r$$
 consists of $r!$ terms.
From (10·12), D can be split into a sum of 2^n determinants, namely,
$$D = t_1 t_2 \dots t_n \left(E + \sum_{i=1}^n E(q_1, q_2, \dots, q_n) e^{r_i t_{g_1} - 1} t_{g_2} - 1 \dots t_{g_r} - 1 \right). \tag{2}$$

$$D = t_1 t_2 \dots t_n \left(E + \sum_{r=1}^{n} {}^*E(g_1, g_2, \dots, g_r) e^r \frac{t_{g_1} - 1}{t_{g_1}} \frac{t_{g_2} - 1}{t_{g_2}} \dots \frac{t_{g_r} - 1}{t_{g_r}} \right), \qquad (10\cdot20)$$
with the abbreviation
$$\sum_{r=1}^{n} {}^* = \sum_{r=1}^{n} {}^*\sum_{g_1, g_2, g_3 = 1}^{n} \dots \frac{t_{g_r} - 1}{t_{g_2}} \dots \frac{t_{g_r} - 1}{t_{g_r}} \right), \qquad (10\cdot21)$$

with the abbreviation
$$\sum_{r=1}^{n} \sum_{r=1}^{n} \sum_{\substack{g_1, g_2, \dots, g_r = 1 \\ g_1 < g_2 < \dots < g_r}}^{n} \sum_{r=1}^{n} \sum_{\substack{g_1, g_2, \dots, g_r = 1 \\ g_1 < g_2 < \dots < g_r}}^{n}$$
(10·21)
Now from (10·1) and (10·19)

Now from (10·1) and (10·19)
$$|E(g_1, g_2, ..., g_r)| e^r \leqslant r! e^r \leqslant (re)^r \leqslant (ne)^r \leqslant {n \choose \sqrt{110}} - 1 r,$$
 (10·22)

Now from (10·1) and (10·19)
$$|E(g_1, g_2, ..., g_r)| e^r \le r! e^r \le (re)^r \le (ne)^r \le {n \choose \sqrt{(\frac{11}{10})}} - 1 \}^r,$$
 hence

hence
$$\left| \sum_{1}^{n} * E(g_1, g_2, ..., g_r) e^r \frac{t_{g_1} - 1}{\frac{t_{g_r} - 1}{t}} ... \frac{t_{g_r} - 1}{\frac{t_{g_r} - 1}{t}} \right| \leqslant \sum_{1}^{n} * \left\{ \sqrt[n]{\left(\frac{11}{10}\right)} - 1 \right\}^r \frac{t_{g_1} - 1}{t} ... \frac{t_{g_r} - 1}{\frac{t_{g_r} - 1}{t}} ... \frac{t_{g_r} - 1}{t} ... \frac{t_{g_r} -$$

$$\left| \sum_{r=1}^{n} {}^*E(g_1,g_2,\ldots,g_r) \, \epsilon^r \frac{t_{g_1}-1}{t_{g_1}} \ldots \frac{t_{g_r}-1}{t_{g_r}} \right| \leq \sum_{r=1}^{n} {}^*\{\sqrt[n]{(\frac{1}{10})}-1\}^r \frac{t_{g_1}-1}{t_{g_1}} \ldots \frac{t_{g_r}-1}{t_{g_r}}$$

$$\left|\sum_{r=1}^n * E(g_1,g_2,\ldots,g_r) \, \epsilon^r \frac{t_{g_1}-1}{t_{g_1}} \ldots \frac{t_{g_r}-1}{t_{g_r}}\right| \leqslant \sum_{r=1}^n * \left\{ \sqrt[n]{\left(\frac{1}{10}\right)} - 1 \right\}^r \frac{t_{g_1}-1}{t_{g_1}} \ldots \frac{t_{g_r}-1}{t_{g_r}}$$

$$\begin{vmatrix} \sum_{r=1}^{n} S(g_1, g_2, ..., g_r) & \frac{1}{t_{g_1}} & \frac{1}{t_{g_r}} \end{vmatrix} \leq \sum_{r=1}^{n} \left(\sqrt{(10)} - 1 \right) \frac{1}{t_{g_1}} & \frac{1}{t_{g_r}} \\ = \prod_{r=1}^{n} \left(1 + \frac{n}{(11)} - 1 \right) \frac{t_g - 1}{t_{g_1}} - 1 - (t_1 t_1 - t_2) \frac{1}{t_{g_1}} \right) \frac{1}{t_{g_1}} + \frac{n}{(11)} \left(t_1 - 1 \right) \frac{1}{t_{g_1}}$$

$$(10)$$

$$=\prod_{n=1}^{n} \left(1 + \left\{\sqrt[n]{\left(\frac{11}{10}\right)} - 1\right\} \frac{t_g - 1}{t}\right) - 1 = (t_1 t_2 \dots t_n)^{-1} \prod_{n=1}^{n} \left\{1 + \sqrt[n]{\left(\frac{11}{10}\right)} (t_g - 1)\right\} - 1, \quad (10)$$

$$=\prod_{g=1}^{n} \left(1 + \left\{\sqrt[n]{\left(\frac{11}{10}\right)} - 1\right\} \frac{t_g - 1}{t_g}\right) - 1 \\ = (t_1 t_2 \dots t_n)^{-1} \prod_{g=1}^{n} \left\{1 + \sqrt[n]{\left(\frac{11}{10}\right)} \left(t_g - 1\right)\right\} - 1, \quad (10 \cdot 23) = 0.$$

$$= \prod_{g=1} \left(1 + \left\{ \sqrt[n]{\left(\frac{1}{10}\right)} - 1 \right\} \frac{\iota_g - 1}{t_g} \right) - 1 = (t_1 t_2 \dots t_n)^{-1} \prod_{g=1}^{n} \left\{ 1 + \sqrt[n]{\left(\frac{1}{10}\right)} \left(t_g - 1 \right) \right\} - 1, \quad (10)$$

$$= \prod_{g=1} \left(1 + \left\{ \sqrt[n]{\left(\frac{1}{10}\right)} - 1 \right\} \frac{\sqrt{g}}{t_g} \right) - 1 = (t_1 t_2 \dots t_n)^{-1} \prod_{g=1} \left\{ 1 + \sqrt[n]{\left(\frac{1}{10}\right)} \left(t_g - 1\right) \right\} - 1, \quad (10)$$

$$-\prod_{g=1}^{n} \left(1 + \sqrt[n]{(\frac{1}{10})} - 1\right) - \frac{1}{t_g} = 1 - (t_1 t_2 \dots t_n) - \prod_{g=1}^{n} \left(1 + \sqrt[n]{(\frac{1}{10})} (t_g - 1)\right) - 1, \quad (10)$$
Thence
$$D \leqslant \prod_{g=1}^{n} \left\{1 + \sqrt[n]{(\frac{1}{10})} (t_g - 1)\right\}, \quad (10)$$

 $D \geqslant 2 \prod_{g=1}^{n} t_g - \prod_{g=1}^{n} \{1 + \sqrt[n]{(\frac{11}{10})} (t_g - 1) \}.$

(10.26)

(10.28)

(10.29)

(10.30)

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 $D < \prod_{g=1}^n \left\{ \sqrt[n]{\left(\frac{1}{10}\right)} + \sqrt[n]{\left(\frac{1}{10}\right)} \left(t_g - 1\right) \right\} = \prod_{g=1}^n \left\{ \sqrt[n]{\left(\frac{1}{10}\right)} \ t_g \right\} < \prod_{g=1}^n \left(\frac{\sqrt[n]{\left(\frac{1}{10}\right)} \ \sqrt[n]{\left(\frac{3}{2}\right)}}{1 - \epsilon} \right) = \frac{3}{20} (1 - \epsilon)^{-n} < \frac{5}{3}.$

Further, since
$$2 - (\frac{11}{10})^{r/n} > 0$$
 for $r = 1, 2, ..., n$, then from (10·16) and (10·25),

$$D \ge 2 \prod_{n=0}^{n} \{1 + (t_{n} - 1)\} - \prod_{n=0}^{n} \{1 + \sqrt[n]{(\frac{11}{10})} (t_{n} - 1)\}$$

 $D\geqslant 2\prod_{g=1}^{n}\left\{1+(t_{g}-1)\right\}-\prod_{g=1}^{n}\left\{1+\sqrt[n]{(\frac{1}{10})}\left(t_{g}-1\right)\right\}$

with D = 1 if and only if $t_1 = t_2 = \dots = t_n = 1$.

 $=1+\sum_{r=1}^{n}*\left\{2-\left(\frac{1}{10}\right)^{r/n}\right\}\left(t_{g_1}-1\right)\ldots\left(t_{g_r}-1\right)\geqslant 1,\quad (10\cdot27)$

This proves that

the lower bound being assumed if and only if $t_1 = t_2 = \dots = t_n = 1$, i.e. if

 $P_1^* = P_1, P_2^* = P_2, ..., P_n^* = P_n.$ Since D > 0, the *n* points $P_1^*, P_2^*, \dots, P_n^*$ are independent; therefore $D = id(\Lambda),$

where j is a positive integer. From (10.9) and (10.28) it follows that $\frac{5}{2} > j \cdot \frac{5}{6}, \quad j < 2,$

and so $j=1, d(\Lambda)=D\geqslant 1$, with equality if and only if $\Lambda=\Lambda_0$. Since Λ_0 is K-admissible and since $d(\Lambda_0) = 1$, this completes the proof that Λ_0 is the only critical lattice

of K, and also that Λ_0 is singular. Corollary. For any given integer $m \ge n$, there exists a bounded star body K with a critical lattice having just 2m points on the boundary of K.

Proof. Nearly obvious, because any star body K' has the required property if it satisfies the following three conditions: (a) K, as defined in the last proof, is a subset of K'. (b) Λ_0 , as defined in the last proof, is K'-admissible. (c) Just 2m points

of Λ_0 lie on the boundary of K'.

provided the boundary consists of a finite number of analytical arcs. This method may be extended to the n-dimensional case, but, naturally, the calculations now become very complicated.

Remark. In an earlier paper on star domains,* I discussed a method by which to obtain $\Delta(K)$ and the critical lattices for every bounded two-dimensional star body,

11. The lattice function $F(\Lambda)$

If Λ is a lattice, t a positive number, and $t\Lambda$ denotes the lattice of all points tP(11.1)

where P runs over Λ , it is obvious that $d(t\Lambda) = t^n d(\Lambda).$ * Mahler—On lattice points in two-dimensional star domains, to appear in the Proceedings of the London Mathematical Society.

K. Mahler 166 Further, if K denotes the star body (not necessarily bounded) of distance function F(X), write $F(\Lambda) = 1.b. F(P),$ (11.2)for the lower bound of F(P) extended over all points $P \neq 0$ of Λ . Then the symbol $F(\Lambda)$ has the following evident properties: Λ is K-admissible if and only if $F(\Lambda) \geqslant 1$. Λ is a critical lattice of K if and only if $F(\Lambda) = 1$, $d(\Lambda) = \Delta(K)$; $F(t\Lambda) = tF(\Lambda)$ if t > 0. further (11.3)A star body is therefore of the finite type if $F(\Lambda) > 0$ for at least one lattice, and is of the infinite type if $F(\Lambda) = 0$ for all lattices. In the special case when K is a bounded star body, it is easily seen that F(A) is a continuous function of Λ ; i.e. if $\Lambda_1, \Lambda_2, \Lambda_3, \dots$ is a convergent sequence of lattices of limit Λ , then $\lim_{r\to\infty} F(\Lambda_r) = F(\Lambda).$ (11.4)If, however, K is an unbounded star body, then $F(\Lambda)$ need not be continuous, as the following example shows. We choose $F(X) = |x_1 x_2 \dots x_n|^{1/n},$ (11.5)and take for Λ the lattice of basis $X_h = (\omega_h^{(1)}, \omega_h^{(2)}, \dots, \omega_h^{(n)}), \text{ where } h = 1, 2, \dots, n,$ (11.6)as defined in the proof of part (1) of theorem 5; there is no restriction in assuming that this basis is reduced. Further, denote by $X_1^{(r)}, X_2^{(r)}, \dots, X_n^{(r)}, \text{ where } r = 1, 2, 3, \dots,$ an infinity of sets of n independent points with rational co-ordinates such that $\lim_{r \to \infty} |X_h^{(r)} - X_h| = 0$, where h = 1, 2, ..., n, (11.7)and such that further $X_1^{(r)}, X_2^{(r)}, \dots, X_n^{(r)}$ form a reduced basis of the lattice A_r generated by these n points. Then by the proof of theorem 5, $F(\Lambda) \geqslant 1$, (11.8) $F(\Lambda_r) = 0$ while, on the other hand, (11.9) $\lim_{r\to\infty} F(\Lambda_r) = 0,$ and (11.10)since a linear form with rational coefficients represents zero. 12. Lattice points near the boundary of an unbounded star body It was seen in § 9 that a critical lattice of any bounded star body has at least 2npoints on its boundary. For unbounded star bodies, this is no longer so; as will be seen in the next paragraph, there exists an unbounded star body of the finite type such that at least one of its critical lattices has no point on its boundary.

 $F\left(\frac{\Lambda}{1+\epsilon}\right) \geqslant 1.$ (12.3)whence Therefore $\frac{\Lambda}{1+\epsilon}$ is also K-admissible, but is of smaller determinant than Λ , and so

A is not critical.

This theorem leads to:

PROBLEM A. Let
$$K: F(X) \leq 1$$
 be a star body of the finite type, A a critical lattice of K , and $\epsilon > 0$ any arbitrarily small number. Do there exist n independent points P_1, P_2, \ldots, P_n

of Λ such that $1 \leqslant F(P_q) < 1 + \epsilon$, where g = 1, 2, ..., n? (12.4)I have not been able to decide this question. The difficulty lies in the fact that

$$F(\Lambda)$$
 may be discontinuous, and so the method of the proof of theorem 11 cannot be applied.
Remark. From theorems 8 and 13, for any given $\epsilon > 0$, every lattice of determinant $d(\Lambda) = \Delta(K)$ contains a point $P \neq O$ satisfying $F(P) < 1 + \epsilon$.

$$l(A) = \Delta(K)$$
 contains a point $P \neq O$ satisfying $F(P) < 1 + \epsilon$.

13. An example of an unbounded star body with no critical

LATTICE POINTS ON ITS BOUNDARY

Theorem 14. Let
$$F_0(X)$$
 be the distance function

THEOREM 14. Let
$$F_0(X)$$
 be the distance function
$$F_0(X) = |x_1 x_2 \dots x_n|^{1/n}, \tag{13.1}$$

nd let further
$$F(X)$$
 be any distance function satisfying the conditions

and let further F(X) be any distance function satisfying the conditions $F(X) \geqslant F_0(X)$ if $F_0(X) > 0$, (13.2)

$$F(X) \geqslant F_0(X) \quad if \quad F_0(X) > 0, \tag{13.2}$$

$$F(X) \geqslant F_0(X) \quad if \quad F_0(X) > 0,$$
 (13.2)
 $\frac{F(X)}{F(X)} \rightarrow 1 \quad if \quad F_0(X) > 0, \mid X \mid^{-1} F_0(X) \rightarrow 0.$ (13.3)

 $\frac{F(X)}{F_0(X)} \rightarrow 1$ if $F_0(X) > 0$, $|X|^{-1}F_0(X) \rightarrow 0$.

$$\frac{F(X)}{F_0(X)} \to 1 \quad if \quad F_0(X) > 0, \mid X \mid^{-1} F_0(X) \to 0. \tag{13}$$

Denote by K_0 and K the star bodies of distance functions $F_0(X)$ and F(X), respectively.

Then

 $\Delta(K) = \Delta(K_0).$ (13.4)

Proof. K is a subset of K_0 , and so from theorem 7, it follows that

 $\Delta(K) \leq \Delta(K_0)$.

(13.5)

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Now assume that	$\Delta(K$	$X) < \Delta$	$(K_0);$		(13.6)
this assumption leads to a cont The function $f(X)$ defined by	radictio	on, as	s will be prove	ed.	
$f(X) = \left\{ \right.$	$\frac{F(X)}{F_0(X)}$ 1	if F	$F_0(X) \neq 0,$ $F_0(X) = 0, X \neq 0$	0,	(13.7)
and not defined if $X = O$, is co unit sphere $ X = 1$. Let $c \ge 1$					s of the
f	$(X) \leqslant c$	if	X =1.		(13.8)
Then, since $f($	f(X) = f	f(X)	for $t \neq 0$,		(13.9)
c is the upper bound of $f(X)$ for	all X =	≠ <i>O</i> , t	herefore		
	F(2)	$(X) \leqslant c$	$F_0(X)$		(13.10)
for all X , since this inequality Let now Λ be any critical lat				·6),	
	d (\angle	1) < <i>1</i>	$ (K_0),$		(13.11)
or, say, d(A) = (1	$(1+\alpha)^{-1}$	$-(n+1)\Delta(K_0),$		(13.12)
where α is some positive numb	er. Put	t			
	(1 +	-α) Λ	$=\Lambda'$,		(13.13)
so that Λ' is $(1+\alpha)$ K-admissible	le, and				
$d(\Lambda')$	= (1+	$(\alpha)^{-1}$	$\Delta(K_0) < \Delta(K_0)$		(13.14)
Denote further by Σ the set is any point of Σ , then	of all p	oints	of Λ' which a	re inner points of H	X_0 . If P
	$(P) \geqslant 1$	$+\alpha$,	$F_0(P) < 1$,		(13.15)
whence	$\frac{F(1)}{F_0(1)}$	$\left \frac{P}{P} \right > 1$	$1 + \alpha$,		(13.16)
and further, from (13·10), F_0	$(P) \geqslant \frac{1}{c}$	F(P)	$\geqslant \frac{1+\alpha}{c} > 0.$		(13.17)
But from (13·3) there exists a	ositive	e num	$_{ m aber}eta$ such th	at	
$\left \frac{F(X)}{F_0(X)} - 1\right \leqslant$	α if	$F_0(X$	$(X) \neq 0, X ^{-1} I$	$V_0(X) < \beta$.	(13.18)
Hence, by the inequalities just	proved	d,			
	P	$^{-1}F_{0}($	$(P) \geqslant \beta$,		(13.19)
and so	P	$\leq \frac{F_0(1)}{h}$	$\frac{(P)}{\beta} < \frac{1}{\beta}$.		(13.20)

On lattice points in n-dimensional star bodies

 $\max(|p_1|, |p_2|, ..., |p_n|) \le |P| < \frac{1}{R};$

 $\mid p_1 \mid \geqslant \mid p_2 \dots p_n \mid^{-1} \left(\frac{1+\alpha}{c}\right)^n > \left(\frac{1+\alpha}{c}\right)^n \beta^{n-1}.$

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(13.21)

(13.22)

(13.23)

(13.29)

(13.30)

Denote by r any positive integer, and by $Y = \Omega_r X$ the unimodular linear transformation

and

and so, finally,

 $y_1 = r^{n-1}x_1, \quad y_2 = r^{-1}x_2, \dots, y_n = r^{-1}x_n.$ (13.24)

Further denote by $A_r = \Omega_r A'$ the lattice of all points $Q = \Omega_r P$ where P runs over Λ' , and by $\Sigma_r = \Omega_r \Sigma$ the set of all points $Q = \Omega_r P$ where P lies in Σ . Then obviously $d(\Lambda_x) = d(\Lambda'),$ (13.25)

and Σ_r consists of all and only all those points of Λ_r which are inner points of K_0 .

If $P = (p_1, p_2, ..., p_n)$ is a point of Σ and $Q = \Omega_r P = (q_1, q_2, ..., q_n)$ is the corresponding point of Σ_r , then, from (13·23)

$$|q_1| = r^{n-1} |p_1| > \left(\frac{1+\alpha}{c}\right)^n (\beta r)^{n-1},$$
 (13.26)

 $|Q| \geqslant |q_1| > \left(\frac{1+\alpha}{c}\right)^n (\beta r)^{n-1}.$

(13.27)and so As in § 8, denote by K_0^t , where t > 0, the set of all points X of K_0 for which $|X| \le t$.

Then the last inequality for Q shows that there exists a monotone increasing

function R(t) of t such that (13.28)

 Λ_r is K_0^t -admissible if $r \ge R(t)$.

Now the sphere $|X| \leq 1$ is obviously a subset of K_0 , hence also of K_0' if $t \geq 1$.

Therefore, from (13.28),

 $\mid Q \mid \geqslant 1$ for all points $Q \neq O$ of A_r if $r \geqslant R(t)$ and $t \geqslant 1$.

Also since

 $d(\Lambda_r) = d(\Lambda')$ for r = 1, 2, 3, ...,

the sequence of lattices $\Lambda_1, \Lambda_2, \Lambda_3, \dots$ is bounded.

But then, by theorem 2, this sequence contains a convergent infinite subsequence

 $\Lambda_r, \Lambda_r, \Lambda_r, \dots,$

of lattices

say of limit Λ^* . Since, from (13·14), $d(\Lambda^*) = \lim_{k \to \infty} d(\Lambda_{r_k}) = d(\Lambda') < \Delta(K_0),$

170 K. Mahler Λ^* cannot be K_0 -admissible; there is then a point P^* of Λ^* which is an inner point of K_0 and so also an inner point of K_0^t if t is sufficiently large. Further, as in earlier proofs, it may be shown that there are points $P_{r_*}, P_{r_*}, P_{r_*}, \dots$ of $\Lambda_{r_*}, \Lambda_{r_*}, \Lambda_{r_*}, \dots$ respectively, $\lim_{k\to\infty} |P_{r_k} - P^*| = 0.$ (13.31)such that But then P_{r_k} is also an inner point of K_0^t if k is sufficiently large, contrary to (13.28). This completes the proof. Theorem 15. There exists an unbounded star body of the finite type with a critical lattice which has no points on the boundary of this body. *Proof.* The same notation is used as in theorem 14, but it is assumed that F(X)satisfies, instead of (13.2), the stronger condition $F(X) > F_0(X)$ if $F_0(X) > 0$; (13.32) $F(X) = F_0(X) \left\{ 1 + \frac{F_0(X)}{|X|} \right\}.$ e.g. take (13.33)Let Λ be a critical lattice of K_0 . Since K is a subset of K_0 , Λ is K-admissible; further, since from theorem 14, $d(\Lambda) = \Delta(K_0) = \Delta(K),$ (13.34) Λ is a critical lattice of K. But the boundary of K consists only of inner points of K_0 , and so no point of Λ may lie on the boundary of K, as asserted. † It is easily proved from §15 that K_0 and so also K have an infinity of critical lattices. The question also arises: Problem B. Do there exist critical lattices of K_0 which are not critical lattices of K, and do these lattices have points on the boundary of K? 14. Star bodies with automorphisms Let $X = (x_1, x_2, ..., x_n)$ and $X' = (x'_1, x'_2, ..., x'_n)$ be two points in R_n . The linear substitution Ω : $x'_g = \sum_{h=1}^n a_{gh} x_h$, where g = 1, 2, ..., n, (14.1) $\omega = |a_{ah}|_{a,h=1,2,...,n} \neq 0,$ of determinant (14.2) $X' = \Omega X$ or shorter (14.3) $X = Q^{-1}X'$ (14.4)has an inverse The substitution defines a one-to-one mapping of R_n on itself. † A much simpler proof of theorem 15 will be given in Part II of this paper.

 $d(\Omega \Lambda) = |\omega| d(\Lambda).$ (14.5)THEOREM 16. Let $K: F(X) \leq 1$ be a star body of the finite type, Ω a substitution of determinant $\omega \neq 0$, F'(X) the distance function (14.6)

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(14.8)

(14.9)

(14.10)

 $F'(X) = F(\Omega X).$ and K' the star body $F'(X) \leq 1$. Then K' is also of the finite type, and

 $\Delta(K') = |\omega|^{-1} \Delta(K).$ (14.7)*Proof.* If Λ is any K-admissible lattice, then $\Lambda' = \Omega^{-1}\Lambda$ is evidently K'-admissible, and so K' is also of the finite type. Further $\Delta(K')$ is not greater than the lower bound of $d(\Omega^{-1}\Lambda) = |\omega|^{-1} d(\Lambda)$ extended over all K-admissible lattices, i.e.

 $\Delta(K') \leq |\omega|^{-1} \Delta(K)$. Since $F(X) = F'(\Omega^{-1}X)$, conversely

 $\Delta(K) \leq |\omega| \Delta(K')$.

From these two inequalities, the assertion follows at once.

Definition 8. The linear substitution $X' = \Omega X$ is called an automorphism of the

star body $K: F(X) \leq 1$, if identically in X, F(X') = F(X).

It is obvious that such an automorphism leaves both K and its boundary C

where P belongs to Λ ; obviously

invariant. THEOREM 17. If the star body K is of the finite type and admits the automorphism $X' = \Omega X$ of determinant ω , then $\omega = \pm 1$.

Proof. By theorem 16, $\Delta(K) = |\omega|^{-1} \Delta(K)$, whence $|\omega| = 1$ since $\Delta(K) \neq 0$.

This theorem shows that star bodies having automorphisms of determinant

 $\omega \neq \pm 1$, are necessarily of the infinite type, e.g. the star body of distance function

 $F(X) = |x_1^2 x_2 \dots x_n|^{1/(n+1)}$ with the automorphism

 $x'_1 = t^{-\frac{1}{2}(n-1)}x_1, \ x'_2 = tx_2, \ \dots, \ x'_n = tx_n \quad (t > 0).$

(14.11)It is obvious that if K is of the finite type, then the set of all automorphisms of K

forms a group. Whether this group is finite or infinite, discrete or continuous, depends on K itself. Definition 9. An unbounded star body K of the finite type is called automorphic if it admits a group Γ of automorphisms Ω with the following property: 'There exists

a positive constant c depending only on K and Γ such that to every point X of K there is an element Ω of Γ satisfying

 $|\Omega X| \leq c$. (14.12)A few examples of automorphic star bodies are given in the next section.

(1) Let $r \ge 0$ and $s \ge 0$ be integers such that r + 2s = n, and let F(X) be the distance function $F(X) = \left| \prod_{\rho=1}^r x_\rho \prod_{\sigma=1}^s (x_{r+\sigma}^2 + x_{r+s+\sigma}^2) \right|^{1/n}.$

15. Examples of automorphic star bodies

It was shown in the first part of the proof of theorem 5 that the star body

$$K: F(X) \le 1$$
 is of the finite type if $r = n, s = 0$. Just the same proof applies when $s > 0$, except that the field \Re there must now be algebraic with r real and $2s$ complex conjugate fields. If the trivial cases $r = 1, s = 0$ and $r = 0, s = 1$ be excluded, then

(15.1)

K is not bounded and admits a continuous group of automorphisms depending on n-1 parameters, namely, the group of substitutions (15.2) $x'_{\rho} = t_{\rho} x_{\rho}$, where $\rho = 1, 2, ..., r$,

$$x'_{\rho} = t_{\rho} x_{\rho}$$
, where $\rho = 1, 2, ..., r$, (15·2)
 $x'_{r+\sigma} = t_{r+\sigma} x_{r+\sigma} - t_{r+s+\sigma} x_{r+s+\sigma}$, $x'_{r+s+\sigma} = t_{r+s+\sigma} x_{r+\sigma} + t_{r+\sigma} x_{r+s+\sigma}$, $\sigma = 1, 2, ..., s$. (15·3)

while $t_1, t_2, ..., t_n$ are n real numbers such that $\prod_{\rho=1}^{r} t_{\rho} \prod_{\sigma=1}^{s} (t_{r+\sigma}^{2} + t_{r+s+\sigma}^{2}) = \pm 1.$

$$\prod_{\rho=1} t_{\rho} \prod_{\sigma=1} (t_{r+\sigma}^2 + t_{r+s+\sigma}^2) = \pm 1.$$
The star body K is automorphic since obviously every point X of K can be transformed into a point X' of bounded distance from Q by one of these automorphisms.

formed into a point X' of bounded distance from O by one of these automorphisms. (2) Let r be an integer such that $1 \le r \le n-1$, and let K be the star body of distance function

$$F(X) = \left| \sum_{\rho=1}^{r} x_{\rho}^2 - \sum_{\sigma=r+1}^{n} x_{\sigma}^2 \right|^{\frac{1}{2}}. \tag{15.5}$$
For the theory of another forms, K admits a group of automorphisms depending

By the theory of quadratic forms, K admits a group of automorphisms depending on $\frac{1}{2}n(n-1)$ real parameters. It is again possible to show that every point in K can

be transformed by one of these automorphisms into a point of bounded distance from O. Hence K is automorphic provided it is of the finite type, and so the following

problem arises:

Problem C. Is the star body of distance function

F(X) =
$$\begin{vmatrix} r \\ r \end{vmatrix}$$
 r^2 =

that K is of the infinite type if $n \ge 5$.

where

 $F(X) = \left| \sum_{\rho=1}^{r} x_{\rho}^{2} - \sum_{\sigma=1}^{n} x_{\sigma}^{2} \right|^{\frac{1}{2}}$ (15.6)

of the finite or of the infinite type?† For $2 \le n \le 4$, K is of the finite type, because there exist indefinite quadratic forms

in n variables with integral coefficients and of given signature which do not repre-

sent zero non-trivially. If, however, $n \ge 5$, then, by Meyer's theorem (Bachmann 1898), every indefinite quadratic form with integral coefficients does represent zero; so the solution of problem C may be difficult. † Addition, May 1946. In a joint paper, H. Davenport and H. Heilbron have just shown

Theorem 19. Let $\Lambda_1, \Lambda_2, \Lambda_3, \dots$ be a convergent sequence of lattices, say of limit Λ ,

and assume that $\phi = \lim F(\Lambda_r)$ exists and is positive. Let there also be a constant c > 0

and an infinite sequence of points P_1, P_2, P_3, \dots such that

 $\lim F(P_r)$ exists and is equal to ϕ .

(16.8)

 $P_r \neq 0$; $|P_r| \leq c$; P_r lies in Λ_r , where r = 1, 2, 3, ...,

K. Mahler 174 $\lim F(\Lambda_r) = F(\Lambda),$ (16.9)Thenand there exists a point $P \neq O$ of Λ such that F(P) = F(A). (16.10)*Proof.* There is a positive number ρ such that the sphere $|X| \leq \rho$ is contained in the star body $F: F(X) \leq 1$. Put (16.11) $\sigma = \frac{1}{2}\rho\phi$. Then the sphere $|X| \le \sigma$ is contained in the star body $F(X) \le \sigma/\rho$, i.e. in $F(X) \le F(\Lambda_r)$, for all sufficiently large r, say for $r \ge r_0$. Therefore for every point $Q \ne 0$ of Λ_r , since $F(Q) \geqslant F(\Lambda_r),$ $|Q| \geqslant \sigma$ if $r \geqslant r_0$. (16.12)Let, in particular, $Y_1^{(r)}, Y_2^{(r)}, ..., Y_n^{(r)}$ be a reduced basis of Λ_r and $Y_1, Y_2, ..., Y_n$ a basis of Λ taken such that $\lim_{r\to\infty} |Y_g^{(r)} - Y_g| = 0, \quad \text{where} \quad g = 1, 2, ..., n.$ (16.13) $|Y_a^{(r)}| \geqslant \sigma$ for $r \geqslant r_0$, g = 1, 2, ..., n. (16.14)Then On the other hand, from theorem 1, $|Y_1^{(r)}| |Y_2^{(r)}| \dots |Y_n^{(r)}| \leq \gamma_n d(\Lambda_r).$ (16.15) $\lim_{r \to \infty} d(\Lambda_r) = d(\Lambda),$ (16.16)Also, from the hypothesis, $\frac{1}{2}d(\Lambda) \leq d(\Lambda_n) \leq 2d(\Lambda)$ for $r \geq r_1$, say, (16.17)hence and so $|Y_{\sigma}^{(r)}| \leq 2\sigma^{-(n-1)}\gamma_n d(\Lambda)$ for $r \geq \max(r_0, r_1)$, where g = 1, 2, ..., n. (16·18)

Since P_r is a point of Λ_r different from O,

 $P_r = u_1^{(r)} Y_1^{(r)} + \dots + u_n^{(r)} Y_n^{(r)}$ (16.19)

with integral coefficients $u_1^{(r)}, \ldots, u_n^{(r)}$ not all zero. On solving this vector equation for $u_1^{(r)}, \ldots, u_n^{(r)},$

 $d(A_r) \mid u_q^{(r)} \mid = \left| \left\{ Y_1^{(r)}, Y_2^{(r)}, \dots, Y_n^{(r)} \right\} \mid \left| u_q^{(r)} \right| = \left| \left\{ Y_1^{(r)}, \dots, Y_{g-1}^{(r)}, P_r, Y_{g+1}^{(r)}, \dots, Y_n^{(r)} \right\} \right|.$

(16.20)

Hence the lower bound for $d(\Lambda_r)$ and the upper bounds for $Y_q^{(r)}$ and P_r imply that

 $|u_a^{(r)}| \leq c'$ (16.21)

where c' is a positive number independent of r and g.

There exists then an infinite sequence of indices

(16.22)

 $r = r_1, r_2, r_3, \dots, \quad \text{where} \quad \lim_{k \to \infty} r_k = \infty,$

such that the coefficients

 $u_q^{(r_k)} = u_q \text{ say}, \text{ where } k = 1, 2, 3, ...; g = 1, 2, ..., n,$

(16.23)

On lattice p	points in n-dimensional star bodies	175
assume integral values indep $u_1,, u_n$ is different from zero.	endent of k , and such that at least one of the ro. Further	ese integers
$P_{r_k} = u_1 Y_1^{(r_k)}$	$+ + u_n Y_n^{(r_k)}$, where $k = 1, 2, 3,$,	(16.24)
and so the points P_{r_k} tend to	the limit point	
	$P = u_1 Y_1 + \ldots + u_n Y_n + O$	(16.25)
which is a point of Λ . From	the hypothesis	
F(P) = K	$\lim_{r\to\infty} F(P_{r_k}) = \lim_{r\to\infty} F(P_r) = \lim_{r\to\infty} F(\Lambda_r),$	(16.26)
whence	$F(\varLambda) \leqslant \lim_{r \to \infty} F(\varLambda_r).$	(16.27)
Moreover, from the last theo	orem,	
	$F(\Lambda) \geqslant \lim_{r \to \infty} F(\Lambda_r),$	$(16 \cdot 28)$
and so the assertion follows		
17. Lattice points on	THE BOUNDARY OF AN AUTOMORPHIC STAF	R BODY
•	$(X) \leqslant 1$ be an automorphic star body, and let Λ be exists a lattice Λ^* and a point P^* of Λ^* such	-
$F(P^*)$	$= F(\Lambda^*) = F(\Lambda), d(\Lambda^*) = d(\Lambda).$	(17.1)
(Remark. Λ^* need not be d but then is nearly trivial.)	ifferent from Λ . The theorem remains valid in	if $F(\Lambda) = 0$,
<i>Proof.</i> Assume that Λ con	ntains no point P such that	
	$F(P) = F(\Lambda);$	$(17 \cdot 2)$
otherwise the assertion is cepoints $P_1, P_2, P_3 \dots$ of Λ such	rtainly true. There exists then an infinite sthat	sequence of
	$ \lim_{r\to\infty} F(P_r) = F(\Lambda) > 0; $	(17.3)
assume that all these points For each point P_r select an	are different from O . In automorphism Ω_r of K such that	
	$\left \varOmega_{r} P_{r} \right \leqslant c.$	$(17 \cdot 4)$
Put	$\Omega_r P_r = Q_r, \Omega_r \Lambda = \Lambda_r,$	(17.5)
so that Q_r belongs to Λ_r , is d	different from O , and satisfies the inequality	
	$\mid Q_r \mid \leqslant c$.	(17.6)
By the invariance of K ,	$F(Q_r) = F(\Omega_r P_r) = F(P_r),$	(17.7)
hence from the hypothesis	$\lim_{r\to\infty} F(Q_r) = F(\Lambda) > 0.$	(17.8)

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Further, from theorem 17, Ω_r is of	determinant ± 1 , and so	
	$d(\Lambda_r) = d(\Lambda).$	(17.9)
Next, it is shown that	$F(\Lambda_r) = F(\Lambda).$	(17.10)
For if P runs over all points of Λ – and vice versa. But by the invaria		points of $\Lambda_r - [O]$,
and by definition,	F(Q) = F(P),	$(17 \cdot 11)$
$F(\varLambda) = \mathop{\mathrm{l.b.}}_{P ext{ in } \varLambda - [O]}$	$F(P)$, $F(\Lambda_r) = \underset{Q \text{ in } \Lambda_r - [O]}{\text{l.b.}} F(Q)$,	(17.12)
whence (17·10) follows at once. Finally, the sequence of lattices	S	
	$\Lambda_1, \Lambda_2, \Lambda_3, \dots$	
is bounded. For from (17.9), the de	eterminants $d(\Lambda_r)$ are bounded, a	and from $(17\cdot10)$,
$F(Q) \geqslant F(A)$) for all points $Q \neq O$ of Λ_r .	$(17 \cdot 13)$
Hence, if ρ is any number such t contains the sphere $ X \leq \rho F(\Lambda)$,		$\leq \rho$, i.e. $F(\Lambda) K$
$ Q \geqslant F(\Lambda) \rho$	o for all points $Q \neq O$ of Λ_r .	$(17 \cdot 14)$
From theorem 2, there exists th	en an infinite subsequence of lat	tices
	$A_{r_1}, A_{r_2}, A_{r_3}, \dots$	
which tends to a limit, say the late	tice Λ^* ; from (17·9)	
$d(\Lambda^*)$	$= \lim_{k \to \infty} d(\Lambda_{r_k}) = d(\Lambda).$	(17.15)
Hence the supposition of theorem sequence of lattices $\{A_r\}$, the lattice the sequence of lattices $\{A_{r_k}\}$, the lattice present proof. The assertion follows	ce Λ , and the sequence of points lattice Λ^* , and the sequence of points we therefore at once from theorem	$\{P_r\}$ respectively, oints $\{Q_{rk}\}$ of the n 19.
Remark. Theorem 20 does not a	assert that every lattice Λ^* satisf	fying
$F(\Lambda^*) =$	$= F(\Lambda), d(\Lambda^*) = d(\Lambda)$	(17.16)
contains a point P^* such that $F(P$ Then, as follows from results in (Koksma 1936), there exists an inf	the theory of indefinite binary	
F(A)	*) = 1, $d(\Lambda^*) = 3$,	$(17 \cdot 17)$
and some, but not all, of these latt	tices contain points P^* such that	;
	$F(P^*) = 1.$	(17.18)
The following particular case of	the last theorem is of special int	erest.

point on the boundary of K. *Proof.* A lattice Λ is a critical lattice of K if and only if

at once from theorem 20.

 $F(\Lambda) = 1, \quad d(\Lambda) = \Delta(K).$ Now, from theorem 8, critical lattices of K do exist; the assertion follows therefore

Theorem 21. Every automorphic star body K has a critical lattice with at least one

Problem D. Does every critical lattice of an automorphic star body K have at least The example in theorem 20 does not answer this question, but makes it probable

(17.19)

(17.20)

(18.3)

that the answer is in the negative.

one point on the boundary of K?

Theorem 20 further suggests the following:

PROBLEM E. To study the set d_F of the values of $d(\Lambda)$ where Λ runs over all lattices Λ satisfying $F(\Lambda) = 1$. The set d_F has a smallest element which is, of course, $\Delta(K)$; this number and the

other elements of the set may be considered as the successive minima of the lattice point problem for the body $K: F(X) \leq 1$. Even in the case $F(X) = |x_1 x_2|^{\frac{1}{2}}, d_F$ is a very complicated set (Koksma 1936), and the same is to be expected for other un-

bounded star bodies. It is then rather surprising that in the case of automorphic star bodies, all these minima are actually attained in the sense that to every element δ of d_F there exists a lattice Λ^* and a point P^* of Λ^* such that $F(P^*) = F(\Lambda^*) = 1, \quad d(\Lambda^*) = \delta.$

morphisms Ω of K. We denote by Σ_{Γ} the set of the points X in R_n which have the

following property: 'There exists a positive number a(X) depending only on X such that

$$|\Omega X| \leq a(X) \text{ for all } \Omega \text{ in } \Gamma.$$
 (18·1)

This set Σ_{Γ} is called the invariant manifold of K. It may contain only the origin, and it has the following properties:

Let $K: F(X) \leq 1$ be an automorphic star body, and let Γ be a group of auto-

(a) If X lies in Σ_{Γ} , and Ω is an element of Γ , then $Y = \Omega X$ also lies in Σ_{Γ} , and we

may take a(Y) = a(X). (18.2)

For let Ω_1 be an arbitrary element of Γ . Then $\Omega_2 = \Omega_1 \Omega$ also belongs to Γ , and so by the definition of a(X),

 $|\Omega, Y| = |\Omega, X| \leq a(X).$

(b) If $X_1, X_2, ..., X_m$ is any number of points of Σ_{Γ} , and if $t_1, t_2, ..., t_m$ are real numbers, then $t_1X_1+t_2X_2+\ldots+t_mX_m$ also lies in \varSigma_\varGamma , and we may take

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$$a(t_1X_1+t_2X_2+\ldots+t_mX_m)=\big|\,t_1\,\big|\,a(X_1)+\big|\,t_2\,\big|\,a(X)+\ldots+\big|\,t_m\,\big|\,a(X_m).\eqno(18\cdot 4)$$
 For if Ω is any element of Γ , then

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 $|\Omega(t_1X_1 + \ldots + t_mX_m)| = |t_1\Omega X_1 + \ldots + t_m\Omega X_m|$ $\leqslant |t_1| |\Omega X_1| + \ldots + |t_m| |\Omega X_m| \leqslant |t_1| a(X_1) + \ldots + |t_m| a(X_m).$ (18.5)From (b), Σ_{Γ} is a linear manifold. Let it be of dimension δ where $0 \le \delta \le n$, and let

 P_1, \ldots, P_{δ} be a set of δ independent points of Σ_{Γ} . Then the points X of Σ_{Γ} may be written as $X = \xi_1 P_1 + \ldots + \xi_{\delta} P_{\delta}$ (18.6)

$$X = \xi_1 P_1 + ... + \xi_{\delta} P_{\delta}$$
 (18·6)
with real coefficients $\xi_1, ..., \xi_{\delta}$; conversely, every such point X belongs to Σ_{δ} . On considering this vector equation as a system of n equations for the n co-ordinates, we find on solving for ξ_{δ} . ξ_{δ} that

we find on solving for $\xi_1, ..., \xi_{\delta}$ that $\max(|\xi_1|,...,|\xi_{\delta}|) \leq \gamma |X|,$ (18.7)

where
$$\gamma$$
 is a positive number depending only on the choice of $P_1, ..., P_{\delta}$.
(c) There exists a positive constant b such that if X is any point of Σ_{Γ} , Ω any element of Γ , and $Y = \Omega X$, then

(18.8)

$$b^{-1}|X|\leqslant |Y|\leqslant b|X|.$$

For let
$$X = \xi_1 P_1 + \ldots + \xi_{\delta} P_{\delta}$$
. Then
$$|Y| = |\xi_1 \Omega P_1 + \ldots + \xi_{\delta} \Omega P_{\delta}| \leq \max(|\xi_1|, \ldots, |\xi_{\delta}|)(|\Omega P_1| + \ldots + |\Omega P_{\delta}|)$$

$$\leqslant \gamma \big|\ X\ \big|\ \{a(P_1)+\ldots+a(P_\delta)\}=b\,\big|\ X\ \big|,\quad (18\cdot 9)$$
 where
$$b=\gamma \{a(P_1)+\ldots+a(P_\delta)\}.$$

Further if X is in Σ_{Γ} and $Y = \Omega X$, then Y is also in Σ_{Γ} and $X = \Omega^{-1}Y$. Hence by

the same proof $|X| \leq b|Y|$, whence the assertion. Let now $J_{\Gamma} = K \times \Sigma_{\Gamma}$ be the set of all points of Σ_{Γ} which belong to K; we call

 J_{Γ} the invariant subset of K.

(d) The invariant subset J_{Γ} is a bounded set. For let X be any point of J_{Γ} . By

definition 9, there exists a positive constant c and an element Ω of Γ such that

 $|\Omega X| \leq c$. (18.10) $|X| \leq b|\Omega X| \leq bc$, Hence from (c), (18.11)

as asserted. This result shows that the dimension δ of Σ_{Γ} and J_{Γ} is at most n-1. For let this

assertion be false so that $\delta = n$. Then Σ_{Γ} coincides with the whole space R_n , and therefore J_{Γ} is identical with K. Hence, from (d), K is a bounded set, contrary to the definition of an automorphic star body.

 $0 \le \delta \le n-2$. Take for K the star body of distance function

$$F(X)=\max{(\{x_1^2+\ldots+x_\delta^2\}^{\frac{1}{2}},\left|\,x_{\delta+1}\ldots x_n\,
ight|^{1/(n-\delta)})},$$
 and for Γ the group of automorphisms

and for
$$\Gamma$$
 the group of automorphisms

$$x_1' = x_1, ..., x_{\delta}' = x_{\delta}, \quad x_{\delta+1}' = t_{\delta+1}x_{\delta+1}, ..., x_n' = t_nx_n,$$

sional linear manifold
$$x_{2} = \dots = x_{n} = 0$$

$$x_{\delta+1}=\ldots=x_n=0.$$
 The automorphic star bodies with $\delta=0$ are of as

$$x_{\delta+1}=\ldots=x_n=0. \tag{18.14}$$
 The automorphic star bodies with $\delta=0$ are of particular interest; then both \varSigma_{\varGamma}

The automorphic star bodies with
$$\delta = 0$$
 are of particular interest; then both Σ_{Γ} and J_{Γ} reduce to the single point O . To this type belong, for instance, all the star bodies considered in §15. In §20, a general property of star bodies with $\delta = 0$ will

bodies considered in § 15. In § 20, a general property of star bodies with
$$\delta=0$$
 will be proved.

THEOREM 22. Let $K: F(X) \leq 1$ be any star body of the finite type. Then there exists to every number $\epsilon > 0$ a positive number $t = t(\epsilon)$ such that every critical lattice Λ of Kcontains at least one point P satisfying the inequalities

ontains at least one point
$$P$$
 satisfying the inequalities
$$1 \leq F(P) < 1 + \epsilon, \quad |P| \leq t. \tag{19.1}$$
 Proof. By the remark to theorem 10, there is a positive number $t^* = t^*(\epsilon)$ such that the star body
$$K^* = K^{(t^*)}: \qquad F(X) \leq 1, \quad |X| \leq t^*$$

Put

then

$$\operatorname{ody}$$

$$t = \left(1 + \frac{\epsilon}{2}\right)t^{s}$$

Put
$$t = \left(1 + \frac{e}{2}\right)t^*, \quad K^{**}$$
 so that K^{**} consists of the points satisfying

$$t = \left(1 + \frac{e}{2}\right)t^*$$

$$t = \left(1 + \frac{\epsilon}{2}\right)t^*, \quad K^{**} = \left(1 + \frac{\epsilon}{2}\right)K^*,$$

$$\Delta(K^*) \geqslant \left(1 + \frac{\epsilon}{2}\right)^{-n} \Delta(K).$$

$$\left(1 + \frac{\epsilon}{2}\right) t^* = K^{**} - \left(1 + \frac{\epsilon}{2}\right)^{-n} \Delta(K).$$

 $F(X) \leqslant 1 + \frac{\epsilon}{2}, \quad |X| \leqslant \left(1 + \frac{\epsilon}{2}\right)t^* = t,$

 $\Delta(K^{**}) = \left(1 + \frac{\epsilon}{2}\right)^n \Delta(K^*) \geqslant \Delta(K).$

 $F(P) \leqslant 1 + \frac{\epsilon}{2} < 1 + \epsilon, \quad |P| \leqslant t;$

Hence every lattice of determinant $\Delta(K)$ contains a point $P \neq 0$ for which

$$1 \le F(P) < 1 + \epsilon$$
, $|P| \le t$.
theorem 10, there is a pos

$$such the valities |P| \leq t.$$

EOR
$$f the$$

; then
$$\Sigma_{arGamma}$$

where
$$t_{\delta+1}, \dots, t_n$$
 are real numbers of product $t_{\delta+1} \dots t_n = 1$; then Σ_{Γ} is the δ -dimensional linear manifold
$$x_{\delta+1} = \dots = x_n = 0. \tag{18.14}$$

the
$$\delta$$
-din

$$(18.14)$$
both Σ_{Γ}
the star
$$= 0 \text{ will}$$

(19.1)

(19.2)

(19.3)

(19.4)

(19.5)

(19.6)

(18.12)

(18.13)

180K. Mahler if the lattice is critical with respect to K, then moreover $F(P) \geqslant 1$. (19.7)whence the assertion. 20. Automorphic star bodies with $\Sigma_{\Gamma} = J_{\Gamma} = \{O\}$ Theorem 23. Let $K: F(X) \leq 1$ be an automorphic star body for which Σ_{Γ} and so also J_{Γ} consist of the single point O. Further let Λ be any critical lattice of K, and ϵ any positive number. Then there exists an infinite sequence of different points P_1, P_2, P_3, \dots of A such that $1 \leqslant F(P_{\mu}) < 1 + \epsilon$, where $\mu = 1, 2, 3, \dots$ (20.1)*Proof.* Assume the assertion is false. There is then a positive number ϵ and a critical lattice Λ of K such that the inequality $1 \leqslant F(P_n) < 1 + \epsilon$ (20.2)is satisfied by only a finite number of points of Λ , say by only the m points $P_1, P_2, ..., P_m;$ by the last theorem, m is not zero. It may be assumed, without loss of generality, that ϵ and Λ have been chosen so as to make m a minimum, that is, There does not exist any positive number e^* and any critical lattice Λ^* of K such that the inequality $1 \leqslant F(P_u^*) < 1 + \epsilon^*$ (20.3)is satisfied by less than m points P_{μ}^* of Λ^* . This minimum assumption implies, in particular, that $F(P_u) = 1$, where $\mu = 1, 2, ..., m$; (20.4)for if, for instance, $F(P_m) = 1 + \delta > 1$, then, on putting $\epsilon^* = \delta$, $\Lambda^* = \Lambda$, there are less than m points P^* of Λ^* such that $1 \le F(P^*) < 1 + e^*$. (20.5)Let now Ω be any automorphism in Γ . Then from (20·2), (20·4) and theorem 22, the lattice $\Omega\Lambda$ has the following properties: There are just m points P^* of $\Omega\Lambda$ for which $1 \le F(P^*) < 1 + \epsilon$. (20.6) $P^* = \Omega P_1, \Omega P_2, \dots, \Omega P_m$ viz. the points (20.7) $F(\Omega P_{\mu}) = 1$, where $\mu = 1, 2, ..., m$. and, in fact, (20.8)There is, moreover, a positive number t independent of Ω and μ such that $|\Omega P_{\mu}| \leq t$ for at least one index μ with $1 \leq \mu \leq m$. $(20.8\frac{1}{2})$ $\Omega_1^{(\mu)} = \Omega_r^{(\mu+1)}, \ \Omega_2^{(\mu)} = \Omega_r^{(\mu+1)}, \ \Omega_3^{(\mu)} = \Omega_r^{(\mu+1)}, \ \dots$

of $\{\Omega_r^{(\mu+1)}\}\$ such that the point sequence $\{\Omega_1^{(\mu)}P_{\mu}, \Omega_2^{(\mu)}P_{\mu}, \Omega_3^{(\mu)}P_{\mu}, \ldots\}$ tends to a limit, say the point P_u^* .

This means that the last sequence $\{\Omega_r^{(1)}\}\$ has the following properties: $\lim_{r\to\infty} |\Omega_r^{(1)} P_m| = \infty.$

If μ is one of the indices 1, 2, ..., m-1, then either $\lim_{r\to\infty} \left| \Omega_r^{(1)} P_\mu \right| = \infty,$

or there exists a finite point P^*_{μ} such that

hence g + h = m. From (20·11) and (20·17), then

Denote then by $\mu_1, \mu_2, \dots, \mu_g$ those different indices μ with $1 \leq \mu \leq m$ for which

by $\mu_1^*, \mu_2^*, \dots, \mu_h^*$ those for which

 $\lim_{r\to\infty} \left| \Omega_r^{(1)} P_\mu - P_\mu^* \right| = 0.$

 $\lim_{r\to\infty} |\Omega_r^{(1)} P_\mu| = \infty,$

 $\lim_{r\to\infty} \left| \Omega_r^{(1)} P_\mu - P_\mu^* \right| = 0;$

 $q \geqslant 1$, $h \geqslant 1$.

(20.17)

(20.18)

(20.19)

(20.20)

(20.21)

(20.22)

(20.15)

(20.16)

Since $\Omega_r^{(1)}$ is an automorphism of K, it is evident that $d(\Omega_r^{(1)}\Lambda) = d(\Lambda) = \Delta(K), \quad F(\Omega_r^{(1)}\Lambda) = F(\Lambda) = 1, \quad \text{where} \quad r = 1, 2, 3, \dots$ (20.23)(for the second equation, compare the proof of theorem 20), and so the lattices

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form a bounded sequence. From theorem 2, one can therefore choose an infinite subsequence $\{\Omega_r\}$ of automorphisms $\Omega_1 = \Omega_{r'}^{(1)}, \, \Omega_2 = \Omega_{r''}^{(1)}, \, \Omega_3 = \Omega_{r'''}^{(1)} \, \dots$ (20.25)

 $\{\Omega_1^{(1)}\Lambda,\Omega_2^{(1)}\Lambda,\Omega_3^{(1)}\Lambda,\ldots\}$

in $\{\Omega_r^{(1)}\}$ such that the corresponding sequence of lattices

 $\Lambda_1 = \Omega_1 \Lambda, \ \Lambda_2 = \Omega_2 \Lambda, \ \Lambda_3 = \Omega_3 \Lambda, \ \dots$ tends to a limit, the lattice Λ^* , say.

Then from (20.23) it follows that

 $\lim_{r \to \infty} d(\Lambda_r) = \Delta(K), \quad \lim_{r \to \infty} F(\Lambda_r) = 1.$

Further, from $(20.8\frac{1}{2})$, and the construction of $\{\Omega_r^{(\mu)}\}\$ and $\{\Lambda_r\}$, each lattice $\Lambda_r = \Omega_r \Lambda$, where r = 1, 2, 3, ...,

 $P^{(r)} = \Omega_r P_{\mu(r)}$ with $1 \leqslant \mu(r) \leqslant m - 1$, contains a point

 $P^{(r)} \neq 0, \quad |P^{(r)}| \leq t, \quad F(P^{(r)}) = 1.$ such that An application of theorem 19 therefore gives

 $d(\Lambda^*) = \lim_{r \to \infty} d(\Lambda_r) = \Delta(K), \quad F(\Lambda^*) = \lim_{r \to \infty} F(\Lambda_r) = 1,$

which means that Λ^* is a critical lattice. Now a consideration analogous to that in earlier proofs makes it evident that the points

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as defined in (20·21), are the only points P^* of A^* such that

dimension δ ?

able to solve:

 $F(P_{\mu,*}^*) = F(P_{\mu,*}^*) = \dots = F(P_{\mu,*}^*) = 1.$ moreover

therefore false and the assertion is true.

 $1 \leqslant F(P^*) < 1 + \frac{\epsilon}{2};$

number h. This contradicts the minimum assumption $(20\cdot3)$; the hypothesis is

Problem F. Does the assertion of theorem 23 remain true if Σ_{Γ} is of positive

Closely related to problem F is the following question which I also have not been

 $P_{\mu_1^*}^*, P_{\mu_2^*}^*, ..., P_{\mu_1^*}^*$

(20.32)(20.33)Hence Λ^* is a lattice of the same type as Λ , except that m is replaced by the smaller

(20.24)

(20.26)

(20.27)

(20.28)

(20.29)

(20.30)

(20.31)

with the following two properties: (a) The invariant manifold Σ_{Γ} is of positive dimension. (b) There exists a critical lattice Λ of K and a positive number α such that $F(P) \geqslant 1 + \alpha$ (20.34)

for all points
$$P$$
 of Λ which do not belong to Σ_{Γ} .

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(21.1)

(21.2)

(21.3)(21.4)(21.5)

(21.6)

(21.7)

The considerations in §18 can be generalized and lead to the following definition:

Definition 10. Let $K: F(X) \leq 1$ be a star body of the finite type with a group Γ of automorphisms Ω , and let δ be an integer such that $1 \leq \delta \leq n-1$. Then K is said to be

21. Star bodies of rank δ

of automorphisms
$$\Omega$$
, and let δ be an integer such that $1 \leq \delta \leq n-1$. Then K is said to be of rank δ with respect to Γ if δ is the largest integer such that to every positive number t^* and to every δ -dimensional linear manifold M containing O there is an element

$$\Omega = \Omega(t^*, M)$$
 of Γ satisfying $|\Omega X| \geqslant t^* F(X)$ for all points X of M .

THEOREM 24. Let K be the star body of distance function
$$F(X) = \begin{bmatrix} r & s & (x^2 + x^2 + x^2) \\ \Gamma & x & \Gamma & (x^2 + x^2 + x^2) \end{bmatrix}^{1/n}, \text{ where}$$

$$F(X) = \left| \prod_{\rho=1}^r x_\rho \prod_{\sigma=1}^s \left(x_{r+\sigma}^2 + x_{r+s+\sigma}^2 \right) \right|^{1/n}, \quad where \quad r+2s = n,$$

$$ho=1$$
 $\sigma=1$ and let Γ be the group of all automorphisms Ω of K defined by

$$x'_{
ho}=t_{
ho}x_{
ho}, \qquad where$$

$$\Omega \colon \begin{cases} x_{\rho}' = t_{\rho} x_{\rho}, & where \quad \rho = 1, 2, ..., r, \\ x_{r+\sigma}' = t_{r+\sigma} x_{r+\sigma} - t_{r+s+\sigma} x_{r+s+\sigma} \\ x_{r+s+\sigma}' = t_{r+s+\sigma} x_{r+\sigma} + t_{r+\sigma} x_{r+s+\sigma} \end{cases} \quad where \quad \sigma = 1, 2, ..., s,$$

$$(x'_{r+s+\sigma} = t_{r+s+\sigma}x_{r+\sigma} + t_{r+\sigma}x_{r+s+\sigma})$$
 where $t_1, t_2, ..., t_n$ are real numbers satisfying

$$\prod_{
ho=1}^{r} t_{
ho} \prod_{\sigma=1}^{s} (t_{r+\sigma}^{2} + t_{r+s+\sigma}^{2}) = 1.$$

$$\begin{array}{ccc}
& & & & & \\
& & & \\
\rho = 1 & \sigma = 1 & \\
\end{array}$$

Further let
$$r \ge 0$$
, $s \ge 0$, $r+s > 1$.

Further let
$$r \ge 0$$
, $s \ge 0$

Further let
$$r\geqslant 0, \quad s\geqslant 0$$
Then K is of rank $r+s-1$ with respect to R

Then
$$K$$
 is of rank $r+s-1$ with respect to $Proof$ An arbitrary linear manifold N

Then
$$K$$
 is of rank $r+s-1$ with respect to R . An arbitrary linear manifold R .

Then
$$K$$
 is of rank $r+s-1$ with respect to Γ .
Proof. An arbitrary linear manifold M

Proof. An arbitrary linear manifold M through O of dimension r+s-1 can be

defined by n-(r+s-1)=s+1 independent homogeneous linear equations $a_{h1}x_1 + a_{h2}x_2 + ... + a_{hn}x_n = 0$, where h = 1, 2, ..., s + 1,

and where the a's are real numbers. Two cases may now be distinguished: (a) First assume that r > 0, and that at least one coefficient

 a_{hk} with $1 \le h \le s+1$, $1 \le k \le r$ is different from zero, say the coefficient a_{11} .

 $a_{11}x_1 + a_{12}x_2 + \ldots + a_{1n}x_n = 0$ (21.8) $x_1 = b_0 x_0 + \ldots + b_n x_n$ for x_1 , (21.9)where $b_2, ..., b_n$ are real numbers; hence there is a positive constant γ such that $|x_1| \leq \gamma \{x_2^2 + \ldots + x_n^2\}^{\frac{1}{2}}$ (21.10)for all points X of M. Put now $t = \gamma^{1/n}t^*$, and apply the automorphism $X' = \Omega X$ defined by

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$$x_1' = t^{-(n-1)}x_1, \ x_2' = tx_2, \ \dots, \ x_n' = tx_n,$$

$$x_1 = t^{n-1}x_1', \ x_2 = t^{-1}x_2', \ \dots, \ x_n = t^{-1}x_n'.$$

$$F(X) = F(X'),$$

and from $(21\cdot10)$ it follows that

collows that
$$|t^{n-1}x_1'|$$

 $\left| t^{n-1}x_1' \right| \leqslant \gamma \left\{ \left(\frac{x_2'}{t} \right)^2 + \ldots + \left(\frac{x_n'}{t} \right)^2 \right\}^{\frac{1}{2}},$

whence

Then, on solving the equation,

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that is

Then

stant γ such that

$$|t^{n-1}x_1'|$$

 $t^n F(X')^n \leqslant \gamma \{x_2'^{\frac{s}{2}} + \ldots + x_n'^{\frac{s}{2}}\}^{\frac{1}{2}} \left| \prod_{\rho=2}^r x_\rho' \prod_{\sigma=1}^s \left(x_{r+\sigma}'^{\frac{s}{2}} + x_{r+s+\sigma}'^{\frac{s}{2}} \right) \right| \leqslant \gamma \{x_2'^{\frac{s}{2}} + \ldots + x_n'^{\frac{s}{2}}\}^{\frac{1}{2}\{1+(r-1)+2s\}}.$

Hence
$$t^nF(X)^n=t^nF(X')^n\leqslant \gamma\big|\,X'\,\big|^n,\quad \big|\,\Omega X\,\big|=\big|\,X'\,\big|\geqslant \gamma^{-1/n}tF(X)=t^*F(X),$$
 as asserted.

(b) Secondly, let either r=0, or assume that r>0, but that all coefficients

vanish.

Then the equations defining M are of the form

Arrange the 2s co-ordinates $x_{r+1}, x_{r+2}, \dots, x_n$ as s pairs

 $a_{h,r+1}x_{r+1} + a_{h,r+2}x_{r+2} + \dots + a_{h,n}x_n = 0$, where $h = 1, 2, \dots, s+1$. (21·17)

that on solving for x_{r+1} , x_{r+s+1} , the following equations are obtained:

 $(x_{r+\sigma}, x_{r+s+\sigma}), \text{ where } \sigma = 1, 2, ..., s.$

Since the s+1 equations defining M are independent, and since there are only s such pairs of co-ordinates, it must be possible to express at least one such pair of these co-ordinates in terms of the others. Now assume this is the pair (x_{r+1}, x_{r+s+1}) , and

 $x_{r+1} = \sum_{\sigma=2}^{s} (b_{\sigma} x_{r+\sigma} + b'_{\sigma} x_{r+s+\sigma}), \quad x_{r+s+1} = \sum_{\sigma=2}^{s} (c_{\sigma} x_{r+\sigma} + c'_{\sigma} x_{r+s+\sigma}),$

where the coefficients b_{σ} , b'_{σ} , c_{σ} , c'_{σ} are real numbers. Hence there is a positive con-

 $x_{r+1}^2 + x_{r+s+1}^2 \le \gamma \sum_{\sigma=2}^{s} (x_{r+\sigma}^2 + x_{r+s+\sigma}^2),$

 a_{hk} with $1 \le h \le s+1$, $1 \le k \le r$

(21.13)(21.14)

(21.15)

(21.11)

(21.12)

(21.16)

(21.18)

(21.19)

(21.20)

 $x'_{r+1} = t^{-(s-1)}x_{r+1}, \quad x'_{r+s+1} = t^{-(s-1)}x_{r+s+1},$

 $x'_{r+\sigma} = tx_{r+\sigma}, \quad x'_{r+s+\sigma} = tx_{r+s+\sigma}, \quad \text{where} \quad \sigma = 2, 3, \dots, s,$

(21.22)

(21.23)

(21.24)

(21.25)

(21.33)

(21.34)

(21.35)

(21.36)

(21.37)

or conversely,
$$x_{\rho} = x'_{\rho}$$
, where $\rho = 1, 2, ..., r$
$$x_{r+1} = t^{s-1}x'_{r+1}, \quad x_{r+s+1} = t^{s-1}x'_{r+s+1},$$

$$x_{r+\sigma} = t^{-1}x'_{r+\sigma}, \quad x_{r+s+\sigma} = t^{-1}x'_{r+s+\sigma}, \quad \text{where} \quad \sigma = 2, 3, ..., s.$$
 Then a prime $x_{r+\sigma} = t^{-1}x'_{r+\sigma}$, $x_{r+s+\sigma} = t^{-1}x'_{r+s+\sigma}$, where $x_{r+\sigma} = t^{-1}x'_{r+s+\sigma}$.

(21.26)F(X) = F(X'),Then again (21.27)and from (21.20)

$$t^{2(s-1)}(x'_{r+1} + x'_{r+s+1}) \leq \gamma t^{-2} \sum_{\sigma=2}^{s} (x'_{r+\sigma} + x'_{r+s+\sigma}) \leq \gamma t^{-2} \mid X' \mid^{2}, \qquad (21 \cdot 28)$$
whence
$$t^{2s} F(X')^{n} \leq \gamma \mid X' \mid^{2} \left| \prod_{\sigma=1}^{r} x'_{\rho} \prod_{\sigma=2}^{s} (x'_{r+\sigma} + x'_{r+s+\sigma}) \right| \leq \gamma \mid X' \mid^{2+r+2(s-1)} = \gamma \mid X' \mid^{n}. \qquad (21 \cdot 29)$$

Hence
$$|\Omega X| = |X'| \geqslant \gamma^{-1/n} t^{2s/n} F(X) = t^* F(X),$$

$$|\Omega X| = |X'| \geqslant \gamma^{-1/n} t^{2s/n} F(X) = t^* F(X),$$

$$(21.30)$$

as asserted. Up to now it has only been proved that the rank δ of K with respect to Γ is at least r+s-1; one now proves that $\delta < r+s$. This is trivial from definition 10 if s=0.

Up to now it has only been proved that the rank
$$\delta$$
 of K with respect to I is at least $r+s-1$; one now proves that $\delta < r+s$. This is trivial from definition 10 if $s=0$. Let therefore $s>0$. Consider the special $(r+s)$ -dimensional linear manifold M_0 defined by the equations

defined by the equations
$$x_{r+s+\sigma}=0, \quad \text{where} \quad \sigma=1,2,...,s.$$
 (21.3)

(21.31)

$$x_{r+s+\sigma}=0, \quad \text{where} \quad \sigma=1,2,...,s.$$
 (21.3) It suffices to prove that, however Ω is chosen in Γ , there is at least one point X

$$x_{r+s+\sigma}=0, \quad \text{where} \quad \sigma=1,2,...,s.$$
 (21.3) It suffices to prove that, however Ω is chosen in Γ , there is at least one point X

It suffices to prove that, however Ω is chosen in Γ , there is at least one point X of

$$w_{r+s+\sigma} = 0$$
, where $0 = 1, 2, ..., 3$. (21.3) t suffices to prove that, however Ω is chosen in Γ , there is at least one point X .

such that
$$|\varOmega X|<\sqrt{(n+1)}\,F(X).$$

t sumces to prove that, nowever
$$\Omega$$
 is chosen in I , there is at least one point X .

$$|QX| < |(n+1)|F(X)| \tag{21.5}$$

 M_0 such that

$$|\Omega X| < \sqrt{(n+1)} F(X). \tag{21}$$

$$|\Omega X| < \sqrt{(n+1)} F(X). \tag{21}$$

(21.32)

$$|\Omega X| < \sqrt{(n+1)} F(X). \tag{21}$$

$$|\Omega X| < \sqrt{(n+1)} F(X). \tag{21}$$
here is no loss of generality in assuming that the point Y is such that

F(X) = 1;

 $\left| \prod_{\rho=1}^r x_\rho \prod_{\sigma=1}^s x_{r+\sigma}^2 \right| = 1,$

Let now Ω be any element of Γ , and X the point above of M_0 . Then the co-ordinates

 $x'_{\rho} = t_{\rho} x_{\rho}$, where $\rho = 1, 2, ..., r$,

 $x'_{r+\sigma} = t_{r+\sigma} x_{r+\sigma}, \quad x'_{r+s+\sigma} = t_{r+s+\sigma} x_{r+\sigma}, \quad \text{where} \quad \sigma = 1, 2, \dots, s,$

 $\prod_{\rho=1}^{r} t_{\rho} \prod_{\sigma=1}^{s} (t_{r+\sigma}^{2} + t_{r+s+\sigma}^{2}) = 1.$

$$|\Omega X| < \sqrt{(n+1)} F(X). \tag{21}$$
 here is no loss of generality in assuming that the point Y is such that

There is no loss of generality in assuming that the point
$$X$$
 is such that

hence the point $X=(x_1,\ldots,x_r,\,x_{r+1},\ldots,\,x_{r+s},\,0,\ldots,0)$ of M_0 satisfies the equation

but is otherwise arbitrary.

of $X' = \Omega X$ take the form

and where

$$\mid arOmega X \mid < \sqrt{(n+1)} \ F(X).$$

$$|arOlimin_{0}| ext{such that} \ |arOlimin_{0}| < \sqrt{(n+1)} \, F(X).$$

186K. Mahler Choose now X in M_0 such that $x_{\rho} = t_{\rho}^{-1}$, where $\rho = 1, 2, ..., r$, (21.38) $\left. \begin{array}{l} x_{r+\sigma} = (t_{r+\sigma}^2 + t_{r+s+\sigma}^2)^{-\frac{1}{2}}, \\ \\ x_{r+s+\sigma} = 0, \end{array} \right\} \quad \text{where} \quad \sigma = 1,2,...,s;$ (21.39)(21.40)then evidently F(X) = 1, as assumed. This choice of X implies that $x'_{\rho} = 1$, where $\rho = 1, 2, ..., r$, (21.41) $x_{r+\sigma}^{2} + x_{r+s+\sigma}^{2} = 1$, where $\sigma = 1, 2, ..., s$, (21.42) $|X'|^2 = r + s < n + 1$. and so (21.43) $|\Omega X| = |X'| < \sqrt{(n+1)} = \sqrt{(n+1)} F(X),$ whence (21.44)as asserted. This completes the proof. Theorem 25. Let $K: F(X) \leq 1$ be a star body of rank δ with respect to Γ , Λ a critical lattice of K, and ϵ an arbitrary positive number. Then there exist $\delta+1$ independent points $P_1, P_2, ..., P_{\delta+1}$ of Λ such that $1 \le F(P_{\mu}) < 1 + \epsilon$, where $\mu = 1, 2, ..., \delta + 1$. (21.45)*Proof.* Let the assertion be false, i.e. assume that there is a critical lattice A_0 of K and a positive number ϵ such that all lattice points P_0 of Λ_0 satisfying $1 \leqslant F(P_0) < 1 + \epsilon$ (21.46)

Proof. Let the assertion be false, i.e. assume that there is a critical lattice
$$A_0$$
 of K and a positive number ϵ such that all lattice points P_0 of A_0 satisfying
$$1 \leqslant F(P_0) < 1 + \epsilon \tag{21.46}$$
 lie in a certain δ -dimensional linear manifold M containing O . From theorem 22, there is a positive number t such that every critical lattice A

of K contains at least one point P such that $1 \le F(P) < 1 + \epsilon$, $|P| \le t$. (21.47)

Further, by the last definition applied with $t^* = t + 1$, there exists an automorphism Ω in Γ such that

 $|\Omega X| \ge (t+1) F(X)$ for all points X in M. (21.48)

 P_1, P_2, P_3, \dots Denote now by

the points of Λ_0 for which

 $1 \le F(P_r) < 1 + \epsilon$, where r = 1, 2, 3, ...;

(21.49)

by hypothesis, these points belong to M. Then the only points Q_r of the lattice

 $\Lambda = \Omega \Lambda_0$ satisfying

 $1 \leq F(Q_r) < 1 + \epsilon$ (21.50)

are those given by $Q_r = \Omega P_r$, where r = 1, 2, 3, ...(21.51)

 $|Q_r| = |\Omega P_r| \geqslant (t+1) F(P_r) \geqslant t+1,$ and for these points (21.52)

contrary to the existence result (21.47). Hence the assertion is true.

On lattice points in n-dimensional star bodies

 $1 \le F(P_q) < 1 + \epsilon$, where g = 1, 2, ..., n,

however small ϵ is chosen. Hence problem A can be solved in this special case, and

I am greatly indebted to Professor Mordell for his help with the manuscript.

From theorems 24 and 25, it is deduced that if K is the star body of distance

(21.53)

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and Λ is any critical lattice of K, then there exist n independent points P_1, P_2, \ldots, P_n of Λ such that

the answer is in the affirmative.

function

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